

1 **Iberian Variscite mining and consumption: distribution and chronological**
2 **framework from Pico Centeno (Encinasola, Spain).**

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23 **Abstract:**

24 This paper focuses firstly on dating directly the exploitation of Pico
25 Centeno's mine 2, by means of AMS-radiocarbon dating combined with
26 OSL dating and profiling; and secondly, indirect dating and
27 establishment of an Iberian wide chronological framework through the
28 analysis of well-dated variscite consumption contexts.

29 The dataset reported in this paper locates the beginning of variscite
30 production at Pico Centeno c. 5200 BC, coinciding in time with alpine
31 jade production or Casa Montero iberian flint production.

32 Variscite consumption is seen as occasional during 5th and 6th millennium
33 BC, and that occurs together with other greenstone. It will be c. 3000 BC
34 when variscite use becomes popular and starts its apogee, appearing
35 almost in every Iberian 3rd millennium BC burial, which is the moment
36 when the alpine jade popularity declines.

37 By the end of this phase, c. 2500 BC, new valuable resources were
38 already in use in the form of signifying items, e.g. asian and african Ivory,
39 Baltic and Sicilian amber or copper-base metallurgy. So the cycle of the
40 variscite starts with the decline jade in the in the 5th – 4th millennium BC,
41 and ends with the appearance of copper, ivory and extra-peninsular
42 amber in the second half of the 3rd millennium BC.

44 **Introduction**

45

46 Archaeological literature devoted to green body ornaments in
47 Prehistoric Europe has thematically focused, almost with exclusiveness,
48 on the quest for an origin of these artefacts.

49 Since the early 20th century, the geographical origin of these '*perles du*
50 *calais*' has moved across continents, from Middle East turquoise mines to
51 European variscite mines, pointing to a French origin firstly at Montebras
52 (Balagny, 1939) or laterly at Pannacé aluminophosphate mines (Forestier
53 et al., 1973a, 1973b; Lheur, 1993; Massé, 1971) and finally to a Spanish
54 origin at Palazuelo de las Cuevas (Arribas et al., 1971, 1970) or Can
55 Tintorer (Alonso et al., 1978; Bosch and Estrada, 1995; Villalba, 2002)
56 variscite mines.

57 From the 1970s, research devoted to '*calaita*' beads has been oriented
58 to locate and characterise new variscite sources. Since then, new source
59 areas have been discovered at Bragança, Northeast Portugal (Meireles
60 et al., 1987), at the Sarrabus deposit, Sardinia (Marini et al., 1989), the
61 variscite and turquoise outcrops of Punta Corveiro in Spain (Moro et al.,
62 1995), and the Pico Centeno variscite mines, also in Spain (Nocete and
63 Linares, 1999) (fig.1). This focus on source location derives from the belief
64 that the relationship between beads and their geological origin is
65 established principally by comparison of the chemical components of
66 the beads and those of the sources (Edo et al., 1995; Dominguez Bella,

67 2004; Odriozola, 2014; Odriozola et al., 2010; Querré et al., 2014, 2008).
68 Increased geochemical analyses of sources was paralleled by increase
69 in numbers of analysed beads, and thus in knowledge of the minerals
70 used in the beadmaking, e.g. green mica, steatite, turquoise, talc,
71 chlorite... -see Vázquez Verela. Thus, callaite and variscite can no longer
72 be considered synonyms, and the validity of traditional analysis of
73 variscite¹ flows and consumption patterns has been brought into
74 question: as the number of analysed beads increases it becomes more
75 obvious that prehistoric communities are using almost any available
76 green mineral for beadmaking, from the Neolithic to the Bronze Age.

77 Present consensus for the geographical origins of European variscite
78 body ornaments points to Palazuelo de las Cuevas (Aliste, Zamora), Can
79 Tintorer (Gavá, Barcelona) and Pico Centeno (Encinasola, Huelva)
80 (Dominguez Bella, 2004; Herbaut and Querré, 2004; Querré et al., 2008,
81 2014; Odriozola et al., 2010; Odriozola, 2014). However, the chronological
82 span of variscite exploitation is uncertain. The chronology of Palazuelo
83 de las Cuevas and Pico Centeno are uncertain. For Palazuelo de las
84 Cuevas Arribas et al. (1970, 1971) proposed the Arabic period based on
85 the assumption that Zamora was the Arab term for emerald. This
86 proposal was rejected by Virgilio Sevillano (1978) and Campano Lorenzo,
87 et al. (1985), who found TSH pottery and prismatic blanks associated with
88 Las Cercas exploitations, but they did not reject previous Prehistoric
89 interpretations (Sanz Mínguez, et al., 1990). More recent research points

¹ Even today most papers devoted to beads still uses callaite and variscite as synonym.

90 to a possible Copper Age exploitation in La Mazada (Esparza Arroyo and
91 Larrazabal Galarza, 2000). Can Tintorer has been extensively dated
92 between c. 4500 – 3500 cal. BC –see table 6.

93

94 <Figure 1: location of known Iberian variscite sources>

95 This chronological uncertainty and the tremendous amounts of non-
96 variscite beads motivated the pioneering work of Jiménez Gómez (1995)
97 at the site of Zambujal, where she attempts to chronologically order
98 variscite consumption pattern based on the radiocarbon dates and
99 mineralogical analysis; concluding that in the second half of the 3rd
100 millennium cal. BC the use of variscite become generalised to the
101 detriment of other greenstones used earlier times. Recently Villalobos
102 García (2012) has shown that this generalisation of variscite use for
103 beadmaking during the second half of third millennium BC also happens
104 in the spanish Meseta Norte.

105 Unfortunately, establishment of variscite flows or consumption patterns
106 has concentrated almost exclusively on source geochemistry and/or
107 consumption, to the detriment of archaeological aspects such as the
108 socioeconomic likelihood for exploitation of the source, and its
109 chronological framework. In consequence, provenance analysis and
110 consumption patterns often pinpoints a geological source that in many
111 cases lacks evidence of exploitation in a relevant timeframe.

112 Unlike settlements, mines usually do not constitute well-stratified sites, but
113 rather a complex system of use, re-use and re-location of products
114 (Frumkin et al., 2014). Little, if any, datable material survives in direct
115 stratigraphic association with the mined surface. In addition, mines
116 usually experience several exploitation events. Therefore, mine dating
117 tends to be complicated and commonly based on 1. time period-
118 specific mine typology and/or mining technology (tool marks and
119 debris); 2. artefacts typology; and 3. indirectly, throughout the dating of
120 contexts of use of mined materials. Generally, dating of a prehistoric
121 mine based on typology and technology is not a straightforward task. In
122 Iberia this has resulted in chronological frameworks that span Late
123 Prehistory without clear delineation into different late prehistoric periods
124 (Domergue, 1990; Hunt, 2003).

125 The chronology of exploitation at Pico Centeno is presently controversial.
126 Pérez Macías (2011, 2008) considers the three trench mines found at Pico
127 Centeno to represent Bronze Age and Roman copper exploitation. In
128 addition, Pérez Macías (2008) argued that certain other evidences of
129 possible prehistoric mining at Pico Centeno instead represent soundings
130 made in 1883 by the «Mina de cobre Santo Tomás» (Jubes and
131 Carbonell, 1920). However, a Neolithic-Chalcolithic variscite exploitation
132 has been proposed recently based on Pico Centeno mine 2 (PCM2)
133 typology, typo-technological marks on the surface of the mine, typology
134 of mining tools, and indirectly, throughout the dating of Pico Centeno

135 bead-worked variscite contexts of use (Odriozola et al., 2010; Odriozola,
136 2014).

137 This paper pursues a double objective, in the one hand, the dating of
138 Pico Centeno's mine 2 (PCM2), and in the other hand, using the absolute
139 dating recorded for PCM2 in conjunction with radiocarbon dates
140 available for well characterised variscite beads, to generate a time
141 model of variscite exploitation (production) and use (consumption
142 pattern).

143 Dating of variscite mines is a crucial step in evaluating variscite
144 production, consumption, and its socio-cultural significance. This paper
145 focuses firstly on dating directly the exploitation of PCM2 by means of
146 combined OSL and with AMS-radiocarbon dating of two singular
147 features, an extraction face in which base level we found charcoal, that
148 we thought were related with firesetting, and a pit which we believed
149 correspond to a later exploitation phase by OSL; and secondly, indirectly
150 dating PCM2 through the analysis of well-dated variscite consumption
151 contexts. Both datasets will be the centre of a debate oriented towards
152 the production of a tentative model for variscite production and
153 consumption at an Iberian framework.

154

155 **PCM2 Direct dating**

156

157 The outcrop of the Pico Centeno aluminophosphate deposit includes

158 three opencast trench mines which typology resembles that of
159 prehistoric exploitation (Shepherd, 1980; Domergue, 1990; Craddock,
160 1995; Hunt, 1996, 2003). In addition to early exploitation, indicated by
161 small cavities oriented in the direction of the variscite veins, which
162 represent extraction, and tool marking scars created by mallets and
163 hammerstones, we also detected excavation by metallic picks that
164 could date from roman times to the 20th century 'copper fever' (Pérez
165 Macías, 2008, after Jubes and Carbonell, 1920).

166 In 2011 we performed an excavation in mine 2 (PCM2), where we
167 conducted 5 test cuts (A-to-E) (fig.2), which recorded features typical of
168 prehistoric activity for this region of Iberia, but also other features
169 suggestive of more recent activity.

170 <Figure 2>

171 Prehistoric features may be summarised by an access ramp cut on the
172 rock, a central transit area and an extraction face with small cavities
173 and concave marks leaved by the impact or rounded edged tools, as
174 the found at PCM2 (Odriozola and Villalobos García 2015) –see
175 Craddock (1995) for a detailed view on mining technology. A 2 meters
176 deep pit nearby the access area, which cut the prehistoric facies, and
177 the marks left in the extraction face by metal tools, above those left
178 behind by stone tools, point to an additional non-prehistoric exploitation.

179 Besides the probability that PCM2 was exploited in recent times, material
180 culture found during excavations, e.g. quartzite cobbles, chisels and

181 picks, points to a prehistoric human activity, most likely related to an
182 early stage of variscite exploitation during Late Prehistory.

183

184 *The stratigraphy*

185

186 Pico Centeno, like many other mines in Iberia, exhibits a long history of
187 use, re-use, and re-location of products and debris, resulting in a
188 complex stratigraphy (fig. 3).

189 <Figure 3: PCM2 stratigraphy>

190 In PCM2, horizontal units were identified through the stratigraphy. These
191 are interpreted as floor units associated with the mining activities and
192 movement of materials to the exterior of the mine. Cutting these floor
193 units, a deep laterally excavated pit was recorded.

194 The last floor unit (test cut A SU 9, test cut B SU 11 and test cut C SU 12/13)
195 used before the abandonment of the mine exhibits numerous stone tools
196 such as picks and wedges showing strong use wear, small production
197 debris with concave marks, and charcoal remains adhering to both to
198 extraction faces and the floor. It is believed that the abundance of
199 charcoal in this base level and on the surface results from the use of fire-
200 setting technology (Craddock 1995; Willies 1994; Weisgerber and Willies
201 2000).

202 Apparently rapid accumulation of large and medium size rock blocks
203 seals these floor units, indicating filling with debris. Over this debris we

204 recorded several apparently slower, natural, accumulation units, which
205 were sealed by the present use-floor.

206 In test cut B was encountered a pit that cuts all units almost from the
207 actual surface level. This pit is filled with several horizontal units (SU 21-to-
208 24). Above these units we recorded several anthropic fast deposition
209 and intentional oblique units that fill the pit with large blocks of stone [SU
210 16-20] (fig. 3 & 4).

211

212 *Luminescence dating*

213 Five samples were taken for quantitative dating analyses at PCM2: three
214 tubes of sediment (3.8 cm diameter, 15 cm length) and a piece of rock
215 were taken from cut B (units 3, 6, 11, 13 & 18 Figure 4; ITNLUM 696-9, Table
216 2), and one tube of sediment was taken from cut D (ITNLUM 701, Table
217 2). Fifteen additional samples were taken for semiquantitative
218 luminescence profiling, using small tubes (2 cm diameter, 5 cm length).
219 Thirteen of these were from the cut B section (Fig 2, Fig 4), and two were
220 taken from the base of the cut C section (Fig 2, Fig 3) (the unexcavated
221 material in cut C was insufficient for a quantitative dating tube).

222 The East section of test cut B was chosen for OSL dating and profiling
223 since it was the larger conserved stratigraphy (fig. 4) and was expected
224 to include both prehistoric facies and the more modern pit fills.
225 Quantitative samples were taken from the least stony layers, plus one
226 rock to test whether the tailings included material to which fire had been

227 set to facilitate mining. The remaining stratigraphic units were sampled
228 for semiquantitative OSL profiling to test the severity of residual signals in
229 the stonier layers and/or help delimit phases of accumulation. The two
230 additional profiling tubes were taken from the remaining regolith at the
231 base of cut C, close to the 14C sample locations to test chronological
232 relationships between the cuts and between luminescence and 14C
233 results. The sample taken for quantitative analysis from the south section
234 of cut D (ITNLUM 701) was designed to help evaluate the chronological
235 relationship between the fills in cut B and layers of tailings spread around
236 the mine site.

237 <Figure 4>

238 <Table 1>

239 *Luminescence and Dosimetric measurements*

240 Luminescence dating measures delayed phosphorescence, obtained
241 from the recombination of electrons and holes in defects of crystalline
242 insulators. Accumulation results from a primary exposure to ionizing
243 irradiation, and recombination of a fraction of those accumulated results
244 from their liberation by secondary irradiation, either by photons (optically
245 stimulated luminescence, OSL) or phonons (thermally stimulated
246 luminescence, TSL) (Burbidge, 2012). The dose of ionizing radiation
247 absorbed by a crystal (Gy) since a liberation event (heating or light
248 exposure), is evaluated by calibrating the OSL or TSL signal obtained from
249 the as-prepared sample material against that produced by controlled

250 irradiations in the laboratory (Burbidge, 2015). Where the accumulated
251 dose of the as-sampled material was produced by ionizing emissions
252 from long-lived radionuclides in a fixed geometry, the dose rate to the
253 crystals in the sample (Gy a^{-1}) may be calculated from parent
254 radionuclide concentrations or direct dose rate measurements of the
255 sample and its environs. From the dose and the dose rate may be
256 calculated the time since the last liberation event (severe heating or
257 light exposure):

258 Luminescence Age (a) = Absorbed Dose (Gy) / Dose Rate (Gy a^{-1}).

259 In the present study sampling and analyses for quantitative
260 luminescence dating followed the approach described in Burbidge et al.
261 (2014), based on a combination of instrumental neutron activation
262 analysis (INAA, Dias et al., 2013; Dias and Prudêncio, 2007; Gouveia and
263 Prudêncio, 2000; Prudêncio et al., 2006), high resolution gamma
264 spectrometry (HRGS, Trindade et al., 2013), field gamma spectrometry
265 (FGS; Trindade et al., 2014), water absorption and retention under free
266 drainage, and optically stimulated luminescence (OSL) measurements.
267 Here, FGS was conducted using both Target Nanospec and HPI Rainbow
268 MCAs with 2"x2" NaI probes.

269 In the laboratory, water content as a fraction of dry (50°C) sample mass
270 was measured as received ("field"; W_f), saturated (W_s), and following
271 free drainage for 1 h (0 days), and 1, 2, 4, and 8 days (W_{D0-8}) (Burbidge et
272 al., 2014). One end of each sediment sample (tube) was sealed with tale

273 and the other closed with nylon mesh, the rock was brushed clean of
274 loose material, and weighted (W_f). Inverted tubes and the rock were
275 soaked overnight in deionised water, and weighed after removal of
276 standing water (W_s). Tubes were drained with seals perforated, on an
277 inclined board, the rock was drained on an inclined sieve (W_{D0-8}). Fills rich
278 in weathered pelitic host rock were highly water retentive, the clast-rich
279 mine-waste and the rock less so. Differences between K and Th
280 concentrations estimated by FGS, HRGS and INAA (Table 2) were
281 accounted for using attenuation by W_f (Burbidge et al., 2014) for deeper
282 fills, and representative drained values for samples from drier layers
283 (ITNLUM 696, 701). Since sampling was conducted in late autumn, in days
284 following rain, the time averaged burial values for fill and rock (ITNLUM
285 696-9) were estimated as the average of each measured in situ water
286 content and that of the driest sample (ITNLUM 701); the maximum value
287 for this was based on its free drainage for 8 days.

288 The mineralisation associated with precipitation of aluminophosphates to
289 produce Variscite at PCM2 resulted also in the presence of moderately
290 elevated levels of Uranium (Table 2). HRGS was thus conducted for each
291 quantitative dating analysis using 23-31 g of milled material, sealed and
292 equilibrated in polystyrene petri dishes, to test for disequilibrium in the
293 upper U-series. 25 emission lines from ^{40}K and the ^{235}U , ^{238}U and ^{232}Th
294 decay series were mass-normalised and compared with the reference
295 samples GSS1, GSS5, GSR6 used for INAA.

296

297 Weighted means over all emissions yielded similar results to INAA (Table
298 2). U dose rates obtained from (unsealed) FGS measurements, after
299 accounting for in situ water content, were 10%-30% lower than
300 INAA(parent ^{238}U) or Wt Mean HRGS (sealed). The FGS result was
301 considered indicative of minimum ^{222}Rn loss in the field, given the wet
302 conditions, and the FGS U values were preferred for age calculations for
303 the fill and mine-waste samples (ITNLUM 696, 697, 699, 701).

304 Cosmic dose rate was estimated by averaging values calculated for as-
305 sampled burial depth in regolith, and the height of the adjacent rock,
306 following Prescott and Stephan (1982), and using a fit to the data of
307 Prescott and Hutton (1988) to include the soft component.

308 <Table 2>

309 The pelitic host rock produced abundant fine silica, agglomerates of
310 which exhibited slow OSL signal decay, poor recycling and strong
311 recuperation in the SAR-OSL protocol (Murray and Wintle, 2000).
312 Relatively small quantities of 90-160 micron quartz were obtained for
313 quantitative OSL estimation of absorbed dose, by repeated
314 disaggregation, sieving, cleaning with HCl, density separation, HF attack
315 (40%, 40 min), and re-sieving. SAR-OSL measurements were made on 3
316 Risø DA-15 and DA-20 readers with integrated $^{90}\text{Sr}/^{90}\text{Y}$ irradiators
317 calibrated relative to the primary ^{60}Co standard of LPSR, CTN metrology

318 laboratory (1, 75±4; 2, 95±3; 3, 111±3, mGys⁻¹). The calibration curve used
319 the following doses: 0(As-prepared), 20, 0, 5, 10, 40, 80, 0, 20, 20(IR) sβ;
320 test dose D_T = 10 sβ for ITNLUM 696, 697, 699, 701, with results fitted using a
321 saturating exponential. Initial tests indicated that the quartz grains from
322 sample PCM3 yielded relatively little signal per unit dose (low OSL
323 sensitivity) but that the accumulated dose of the as-prepared sample
324 was relatively large. Thus, the analysis ITNLUM 698 used the following
325 radiation exposures: 0 (As-prepared), 200, 0, 800, 1600, 3200, 6400, 0, 200,
326 200 (IR); D_T = 50 sβ for, with results fitted using a saturating exponential
327 plus linear function. Preheats were made at 180, 200, 220, 240, 260 and
328 280 °C/30 s to test for differences in absorbed dose estimates as a
329 function of the relative filling of, and/or transfer of charges between,
330 electron and hole traps during irradiation with the calibration and test
331 doses and their OSL measurement. All measurements using the maximum
332 and minimum preheats were rejected in the analyses ITNLUM 696, 697,
333 699, 701, since systematic deviations or increased scatter in absorbed
334 dose values were often observed. For the analysis ITNLUM 698, scatter in
335 absorbed dose estimates was great for all preheats, and the test-
336 normalised OSL signal from the as-prepared material did not intercept
337 the calibration curve in two cases (Table 2). For accepted
338 measurements, average recycling ratios ranged from 0.93 to 1.01, and
339 from 0.96 to 1.00 (0.81, ITNLUM 698) after exposure to infrared light. Zero
340 dose response was less than 6% of absorbed dose. OSL sensitivity was
341 relatively low for samples from the pit (Table 2), and for ITNLUM 698

342 reduced by 50% during the measurement sequence. Signal integrals
343 using the majority of the initial OSL decay, and 'late' background
344 subtraction, were used in all cases: use of the initial signal gradient
345 (second two OSL channels subtracted from the first two channels)
346 resulted in highly dispersed datasets. The weighted mean of the
347 accepted absorbed dose measurements, weighted to inverse variance,
348 appeared representative of the main grouping in the datasets obtained
349 for ITNLUM 696, 697, 699, and 701, and so was used to calculate the
350 central absorbed dose estimates for use in age calculation for all
351 samples. However, the rock (ITNLUM 698) exhibited scatter to very high
352 values of absorbed dose so that the weighted mean is a minimum value
353 only. The results from this sample indicate that the rock was unheated
354 but large finite results were obtained due to poor behaviour within the
355 SAR protocol.

356 For age estimation, measurements were combined following the
357 approach outlined in Burbidge et al. (2014). Alpha, beta and gamma
358 dose rates from the environment surrounding the sample location were
359 estimated from FGS measurements, after correcting for measurement
360 geometry and the difference between in situ and time averaged water
361 contents. Attenuated environmental dose rates were combined with the
362 self dose rate of a volume of representative size and density, based on
363 the holes excavated for FGS measurement, with values estimated by
364 INAA and HRGS (U from FGS) on material from the adjunct sample taken

365 from those holes, and corrected for time averaged water content
366 estimated from the tube samples. Attenuated dose rates from the
367 environment and adjunct were then combined with similarly estimated
368 values for the samples themselves, which were assumed equal to the
369 adjunct in the case of the tube samples. Dose rates to the etched cores
370 of the grains were calculated from these values, and all combined with
371 calculated cosmic dose rates to estimate total dose rates to the grains
372 measured by luminescence. The absorbed dose from the grains, was
373 divided by the dose rate to estimate time since the event of interest
374 (age), and converted to an estimate of calendar date.

375 For luminescence profiling (Sanderson et al., 2001), a basic series of
376 preparatory treatments was used, to separate coarse (90–250 µm) 40%
377 HF etched fractions, and measured using a simple multi-stimulation
378 protocol (Rodrigues et al., 2013; Odriozola et al., 2014). The profiling
379 approaches referred to above have calibrated signals from the as-
380 prepared material on an aliquot to aliquot basis, using signals resulting
381 from a single regenerative dose in the quasi-linear region of the sample's
382 dose response. In the present case the as-prepared signals from the
383 profiling samples varied strongly and were often much greater than
384 those produced by the calibration dose. Extrapolation of the quasi-linear
385 calibration dose response would tend to overestimate signal resulting
386 from doses in the range 10-200 Gy, where effects of saturation in the
387 dose response, particularly from quartz, would be expected to be

388 evident. To help account for signal saturation and permit comparison
389 between all results, absorbed dose estimated from the profiling samples
390 were estimated using a common saturating exponential dose response
391 characteristic (DRC) obtained directly from the profiling measurements.
392 This DRC was defined using the average of the standardised (Roberts
393 and Duller, 2004) post-IR OSL responses to $50 \text{ s}\beta$ (i.e. $D_T I(D=50\text{s}\beta)/I_T$),
394 assuming a single saturating exponential dose response, $I = I_\infty(1-\exp(-D/\bar{D})$
395), with signal at saturation (I_∞) equal to the average dose of signal
396 saturation (\bar{D}), which was found to be 40 Gy, note that D_T in all cases was
397 $10\text{s}\beta$ (Burbidge et al., 2006; Burbidge, 2015). Semiquantitative age values
398 were estimated by interpolating total dose rates of the quantitative
399 dating samples.

400

401 AMS-Radiocarbon dating

402

403 Five radiocarbon measurements were performed using a 1 MV AMS over
404 charcoal sample recovered at the base level of test cut C (unit 12/13).
405 Chemical preparation of the samples followed standard procedures
406 (Santos Arevalo et al., 2009). Soxhlet extraction was applied using
407 hexane, acetone and ethanol before treating the samples with the
408 Acid-Alkali-Acid cleaning procedure. For the AAA procedure HCl 0.5 M
409 and NaOH 0.1 M were used, and time was carefully controlled to avoid
410 severe losses by dissolution.

411 Between 7–10 mg of clean and dry charcoal were combusted at 950 °C
412 for 3 h in a vacuum-sealed quartz tube with CuO and Ag powder.
413 Quartz tubes had been previously baked at 950 °C to eliminate possible
414 organic matter. Produced CO₂ was then reduced to graphite by adding
415 excess H₂ and using cobalt as catalyst. The resultant mixture of graphite
416 and cobalt was pressed into aluminium cathodes and kept on vacuum
417 until measurement (Santos Arevalo et al., 2009).

418 The PCM2 site was very poor in organic materials suitable for
419 radiocarbon dating, and no short lived materials were recovered from
420 excavations: only charcoals from cut C level 12/13 were of sufficient size.
421 AMS-¹⁴C dates are reported in conventional radiocarbon years before
422 present in accordance with (Stuiver and Polach, 1977). The results have
423 been calibrated using Calib 7.1 with intcal'13 (Reimer, 2013). Calibrated
424 age is presented in calendar years cal. BC (2 sigma), and also as BP
425 (table 3).

426 <Table 3>

427 Anthracological analysis was performed on charcoal fragments, after
428 removal of the samples for AMS-radiocarbon dating. Small size and poor
429 conservation of charcoal samples limited the anthracological
430 determinations, so that in some cases it only allowed determination of
431 angiosperm, and in others to propose a taxon (table 3).

432

433 **Indirect dating**

434

435 Ideally evidence should be combined from excavations of several well-
436 dated sites, including beads, pendants or charm-assemblages made
437 from variscite. Unfortunately, accurate identification of bead mineralogy
438 is lacking for most sites in Iberia, where many green minerals other than
439 variscite were used for bead making. In order to realistically define a
440 variscite consumption chronological framework, we need first to identify
441 beads mineralogy.

442 The mineralogical classification of beads by means of portable
443 analytical devices is not a straightforward task and deserves a full-length
444 paper of its own. However, we have developed a simplified approach
445 for the purpose of helping evaluate bead chronologies.

446 The mineralogical identifications of beads in this paper (1392 samples
447 from 42 different sites along the Iberian Meseta and Atlantic Façade) are
448 based on the chemical composition measured by an Oxford Instrument
449 XMET-7500 portable x-ray spectrometer with a Rh tube, a silicon drift
450 detector (SDD), and an automatic 5-position filter changer.

451 Aluminophosphate identification is almost a straightforward task based
452 on Al-to-P atomic ratios. Almost anytime we detect Al-to-P atomic ratio
453 in the compositional range of variscite $[[\text{MPO}_4 \cdot 2\text{H}_2\text{O}]]$, where M = Al^{3+} ,
454 Fe^{3+} , Cr^{3+} , V^{3+} (Larsen, 1942)], from c. 1 to 1.8 (see Odriozola et al., 2010,
455 and Odriozola, 2014), x-ray diffraction confirms that the analysed
456 sample, either geological or beadworked, has variscite as its main

457 crystallographic phase. Thus, here we use the P-to-Al atomic ratio as an
458 indicator to determine variscite as the raw material of beads.
459 Nevertheless, turquoise, crandallite or aheylite may occur separately or
460 as minor crystallographic phases together with variscite (Larsen, 1942). In
461 these cases Al-to-P atomic ratios together with Ca, Cu and Fe values are
462 taken into account to differentiate between minerals.

463 A much more complicated task is that of differentiating green stones
464 formed by sheet silicates (micas, talc-steatite, chlorite or serpentine). We
465 conservatively classified beads by aluminophosphate, K-aluminosilicates,
466 Mg-aluminosilicates, Mg-silicates and other silicates.

467 This methodology appears adequate for the purposes of the present
468 work, where the aim is to evaluate whether green beads are worked out
469 of variscite or other greenstone.

470

471 **Results**

472

473 *Direct dating*

474

475 The AMS 14C results indicate a palimpsest of organic remains at the base
476 of cut C in the Neolithic (CNA 2144, 2148), Iron Age (CNA 2145), and
477 Mediaeval periods (CNA 2146 2147). The quantitative OSL results from fills
478 indicate accumulation of deposits in cut D and lower cut B in the late
479 18th – early 19th century AD (ITNLUM 696 697), and in the late 19th century

480 in upper cut B (ITNLUM 699 701) (Table 2). In the present study profiling
481 was conducted only in the laboratory and so was not able to inform the
482 sampling strategy (c.f. Burbidge et al., 2008; Sanderson and Murphy,
483 2009), but instead was applied in parallel with quantitative dating, to aid
484 interpretation of fill phases and the intensity of accumulation
485 mechanisms that operated during infill of the pit, and to evaluate
486 relationships between sediment layers from around the site that were not
487 amenable to sampling for quantitative analysis (e.g. Odriozola et al.,
488 2014). A frequency plot of the profiling results (f_p , Figure 5) shows how the
489 mineral grains in most of the sampled layers were last exposed to light
490 ca. 7 ka (Neolithic), 0.6 ka (late Mediaeval), and 0.12 ka (Post
491 Mediaeval). These phases correspond with quantitative results obtained
492 by OSL or AMS14C. Many profiling samples, apparently out of
493 stratigraphic sequence, thus appear to contain redeposited material
494 from earlier phases of activity at the site, in which the OSL signal has not
495 been reset: such records are considered useful for interpreting the history
496 of activity at the a site even when removed from their original context, in
497 the same way as survey or recovery of redeposited potsherds (e.g.
498 Deckers et al., 2005; Burbidge et al., 2014).

499 The OSL profiling results (Figure 5) through the stratigraphy of cut B exhibit
500 two phases of elevated semiquantitative age values, from samples
501 taken in the more stony layers (120-70 cm and 35-25 cm). These are
502 considered indicative of redeposition without (complete) liberation of
503 the trapped charge that is the source of the OSL signal, and hence of

504 rapid redeposition. Quantitative and semiquantitative results from less
505 stony layers associated with them, either immediately below or inter-
506 leaved at approximately the same depth, indicate that their
507 accumulation in their present location corresponds with early and late
508 19th century infill phases. Results from the first phase (120-70 cm) are
509 scattered, e.g. the difference between the results from the two aliquots
510 measured per sample is large relative to the geometric average value
511 indicated by the trend line in Figure 5, but many results approximate the
512 reproducible (tightly grouped pair) from the sample at the base of cut B
513 (145 cm). This profiling sample, and one of the samples from cut C, yield
514 indications of late Mediaeval accumulation contemporaneous with the
515 youngest AMS14C result from cut C (CNA 2147), and the other profiling
516 sample from cut C gave results consistent with the oldest, Neolithic, AMS
517 14C results (CNA 2144, 2148).

518 The late 19th century OSL results are thought to relate to soundings made
519 in 1883 by the Mina de Cobre Santo Tomás, during the 'copper fever'²
520 (Pérez Macías, 2008, after Jubes and Carbonell, 1920). However, fills in
521 the lower layers in cut B, and the tailings in cut D unit 3, both relate to an
522 earlier phase of accumulation.

523 The mixture of charcoals from different periods in the same unit, base
524 level (12/13), indicates phases of repeated use of the cut C space, up to
525 the Mediaeval period, and this is also reflected by semiquantitative OSL
526 results from the bases of cut C and cut B. The space surrounding the void

² Copper fever started in Huelva in the mid 19th century.

527 created by the first miners was last used ca. 600 years ago (bonfires at
528 the border of the pit). PCM2 is in the middle of a copper mining belt of
529 Ossa Morena, exploited since Copper Age and reopened at the end of
530 19th century with the arrival of British enterprises on the search for copper.

531 In spite of the calibrated AMS-14C date spam points that PCM2 was
532 probably in use since the transition 6th /5th millennium cal BC (CNA-2141,
533 CNA-2148), which is corroborated by semiquantitative OSL of a small
534 sample of heated sediment, historical occupations are also recorded.

535 Generally speaking, variscite was exploited for beadmaking during Late
536 Prehistory (Villalba et al., 2001) and during Roman times, for both
537 beadmaking and tessellae production (Gutiérrez Pérez et al., 2015).

538 Some weak evidences point to an occasional use during the Bronze Age
539 (Schubart et al., 2004) and in the 18th Century (García-Guinea et al.,
540 2000). It is unlikely that this resource was exploited during 14th century.

541 Pico Centeno is most likely to have been exploited during the Neolithic
542 and Roman periods, when the archaeological record indicates that
543 variscite was used most intensively for beadmaking.

544 <Figure 5>

545

546 *Indirect dating*

547

548 We have built a new dataset of dated sites, together with a bead
549 mineralogical classification (table 4), which creates an ideal framework

550 for the analysis of variscite consumption during Late Prehistory across
551 Iberia. However, we should bear in mind that most of the sites are burials,
552 some of them had long occupations, and accounts for limited number
553 of short-lived 14C dates. If the direct dating of mine contexts has
554 limitations inherent to a continued use of the space, the dating of bead
555 burial contexts (necropolis or settlements) is subject to difficulties of
556 association. Most of the inventoried variscite adornments were collected
557 in early archaeological excavations, without well-defined stratigraphic
558 contexts. However, some trends in bead consumption can be identified.

559 <Table 4>

560 If we plot the calibrated dates together with the mineralogical diversity
561 of bead assemblages for each site (fig. 6) we can observe how variscite
562 use could begin in the 5th millennium BC. In figure 6, it can also be
563 appreciated that variscite is not the main mineral used along the 4th
564 millennium BC and that it is during 3rd millennium BC when the use of
565 variscite became generalised, just as has previously been stated for
566 Portuguese Estremadura (Jiménez Gómez, 1995) and the northern
567 Spanish Meseta (Villalobos García, 2012). By the end of 3rd millennium,
568 the use of variscite suddenly stops towards the use of green micas.

569 <Figure 6>

570

571 **Discussion**

572 Recorded dates for Pico Centeno's PCM2 variscite mine together with
573 typological and technological criteria show strong evidences of
574 intermittent exploitation during specific periods during Late Prehistory,
575 Iron Age, Medieval Age and Modern Era.

576 For the Iberian Peninsula, a generalised use of variscite is known for the
577 beginning of Late Prehistory (Villalba et al., 2001) and for roman times
578 (Gutiérrez Pérez et al., 2015); and its use was sporadic in modern times
579 (Garcia-Guinea et al., 2000). It is, therefore, in the light of these data and
580 the historical importance of metal mining at Southwest Iberia, specifically
581 at Encinasola and its surroundings (Pérez Macías, 2008), that 3rd century
582 BC and 9-10th, 14th, 18-19th centuries AD mining activity at Pico Centeno
583 should be most likely related to the exploration of metallic resources
584 rather than variscite itself. We will, thus, focus on the prehistoric use of
585 variscite that is where our main interest rests.

586 The oldest dates recorded at Pico Centeno for Late Prehistory (5300-5000
587 and 4900-4700 cal. BC) points to an Early and Middle Neolithic
588 exploitation. However, charcoal radiocarbon dates may be biased
589 towards excessive antiquity by old wood (Schiffer, 1986). The lack of
590 recorded settlements from late 6th millennium BC in the surroundings of
591 Pico Centeno would, to some degree support an old wood effect on the
592 dates. Conversely, a closer look at the available chronologies of Iberian
593 sites that undoubtedly accounts for variscite beads would support an
594 early exploitation of variscite, e.g. Cueva del Moro and Cueva de
595 Chaves (Baldellou et al., 2012) and probably Gruta do Caldeirão (Real,

596 1992) during 6th millennium cal. BC and Chousa Nova (Dominguez-Bella
597 and Bóveda, 2011), Dolmen de Alberite (Sttip and Tamers, 1996), or
598 Fuentepecina II (Rojo Guerra et al. 1996; Delibes de Castro and Rojo
599 Guerra, 1997) during 5th millennium cal. BC³.

600 Coeval in time to earliest variscite use and exploitation, according to
601 PCM2 radiocarbon dates, would be the quarrying and spread of green
602 alpine jade axe around Western Europe (Pétrequin et al., 2006). This
603 phenomenon would be fully coincident in the timing of the overall
604 process of discovery, exploitation and network exchange, to that of
605 variscite being perhaps the scale the biggest difference between both
606 phenomena. Coeval in time to alpine jade axe widest spread and
607 exchange maximum intensification, c. 4500 BC (Pétrequin et al., 2006)
608 (Fig. 6 & Tab. 5).

609 <Figure 6>

610 It will be from 4500 cal. BC onwards –mainly in the 4th and first half of 3rd
611 millennia BC- that variscite mining activity increases drastically (Blasco et
612 al., 1992) and the use of variscite become conspicuous in the Iberian
613 Peninsula (Blasco et al., 1997; Bueno Ramírez et al., 2005; Costa et al.,
614 2011; Gonçalves and Reis, 1982; Guitán Rivera and Vázquez-Varela,
615 1975; Villalobos García, 2012). Therefore, the use of variscite beads
616 become extremely popular, reaching a period of maximum widespread

³ The later dates should be interpreted with caution, considering the sample types and contexts

617 and use during the first half of the 3rd millennium BC, while the use of
618 other green stones becomes testimonial (Fig. 7 & Tab. 6).

619 <Figure 7>

620 From c. 2500 BC onwards the situation became the opposite, and
621 variscite use starts to decline coinciding with the generalisation of new
622 valuable resources, e.g. Asian and African Ivory (Schuhmacher et al.,
623 2009; Schuhmacher, 2012), Baltic and Sicilian amber (Murillo-Barroso and
624 Martinón-Torres, 2012) or copper-based metallurgy (Murillo-Barroso and
625 Montero Ruiz, 2012).

626

627 **Concluding remarks**

628

629 AMS-Radiocarbon and OSL dating of PCM2 dataset reported in this
630 paper show a long-term history of use, from the end of 6th millennium cal.
631 BC Neolithic exploitation of variscite to 19th century AD copper soundings
632 made by Mina de Cobre Santo Tomás.

633 OSL dating indicates that the fills and tailings currently evident on site
634 accumulated in their present positions in the late 18th and late 19th
635 centuries. AMS-14C dating of apparently *in situ* material set at the rear of
636 the mine excavation indicated a palimpsest from the Neolithic, Iron Age
637 and Mediaeval periods. Semiquantitative OSL profiling results, from small
638 samples obtained from stony layers and the remnants of excavated fills
639 not amenable to sampling for fully quantitative OSL analysis, corroborate

640 chronological indications from both quantitative OSL and from AMS-14C.
641 Thus, radiocarbon and OSL provide complementary information on
642 different phases of site usage, linked by luminescence profiling.
643 The new dataset reported in this paper support a prolonged, intermittent
644 and low intensity mining activity⁴. It locates the beginning of variscite
645 consumption starts coeval in time to the decline jade in the 5th – 4th
646 millennium cal. BC, and the end coinciding with the appearance of
647 other signifying items in the second half of the 3rd millennium BC, e.g.
648 copper, ivory and extra-peninsular amber. Variscite consumption
649 reaches its apogee c. 3000 cal. BC, when appears almost in every
650 Iberian burial (Jiménez Goméz, 1995; Villalobos García, 2012).,

651

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653

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659 **8. References**

660

⁴ Similar to that of 3rd millennium cal. BC North Iberian copper mining, calculated to be 35 days per annum along 8 centuries (Blas Cortina, 2011)

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1129 Books.
- 1130
- 1131 **Caption to figures**
- 1132
- 1133 Figure 1: location of known Iberian variscite sources. Pico Centeno
1134 coordinates: Lon: 22°51'29"; Lat:5°13'15" (ETRS89).
- 1135 Figure 2: test cut position and 3D model
- 1136 Figure 3: a) PCM2 planimetry and cuts A, B and C North profiles and cut

1137 B East profile stratigraphies; b) Harris matrix of Cut B (software matrix harris
1138 composer).

1139 Figure 4: a) cut B East profile stratigraphy and b) cut D North profile
1140 stratigraphy, with detailed quantitative OSL sample position (big dots)
1141 and semiquantitative OSL profiling (small dots).

1142 Figure 5: Chronological results from PCM2; note log time axis

1143 Figure 6: radiocarbon calibrated age sum probability plot and pie chart
1144 mineralogical identifications for studied sites.

1145 Figure 7: calibrated data for european flint, jade, variscite and copper
1146 mining. With detail calibrated multisample probability plot of PCM2
1147 radiocarbon dates and sum probabiolity plot flint, jade, variscite and
1148 copper european radiocarbon dates.

1149

1150 **Tables:**

1151

1152 Table 1. Brief description of the OSL dating sampled units.

1153 Table 2a: Luminescence dating measurements: radionuclide
1154 concentrations and water content estimates.

1155 Table 2b: Luminescence dating measurements: summary dose rate,
1156 dose, and calendar date estimates.

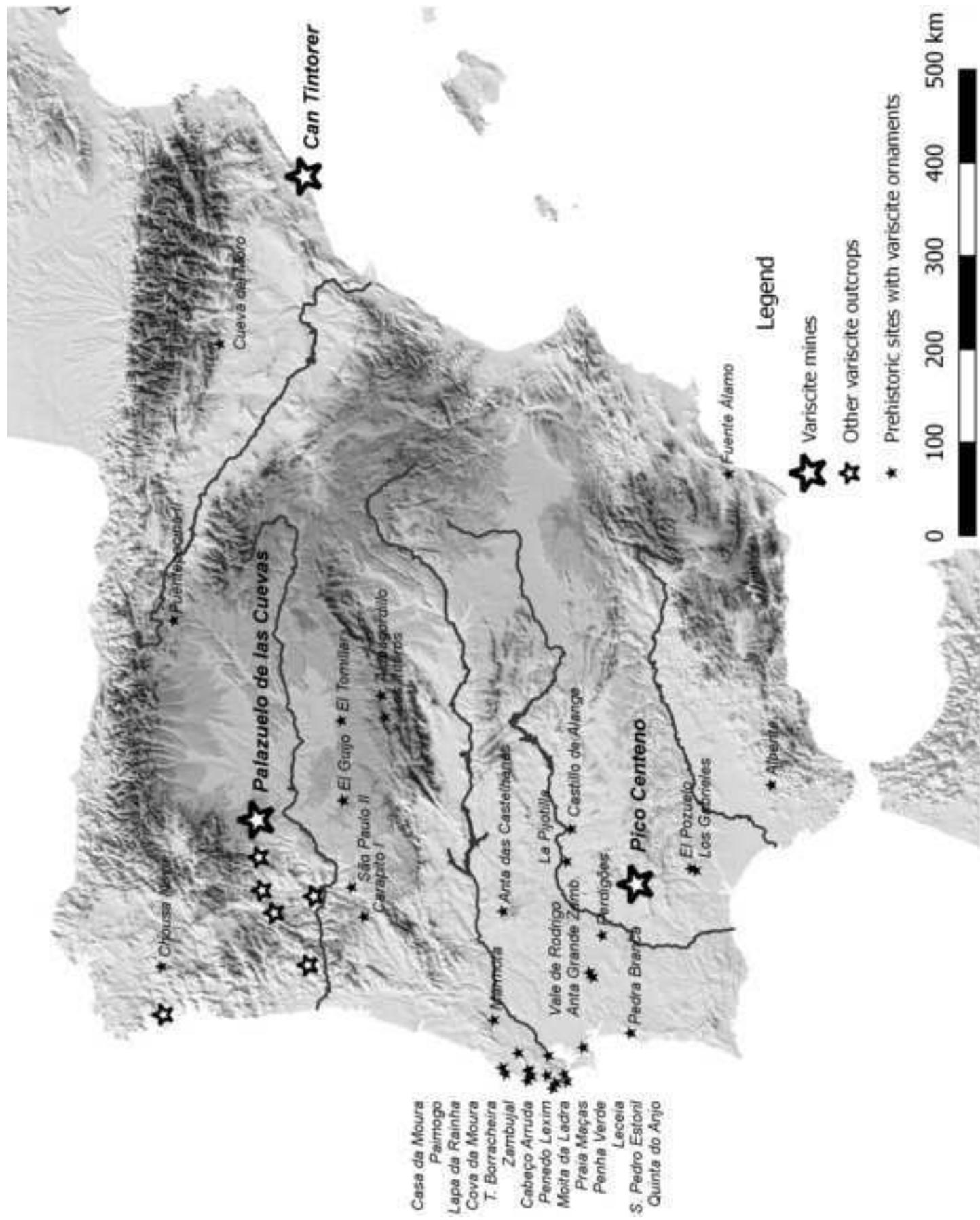
1157 Table 3. Results of AMS-radiocarbon dating and anthracological analysis
1158 of sampled charcoals from PCM2.

1159 Table 4. Results of AMS-radiocarbon dating and anthracological analysis
1160 of sampled charcoals from PCM2.

1161 Table 5. European mining resources available radiocarbon dates

1162 Table 6: Studied sites available radiocarbon dates

1163



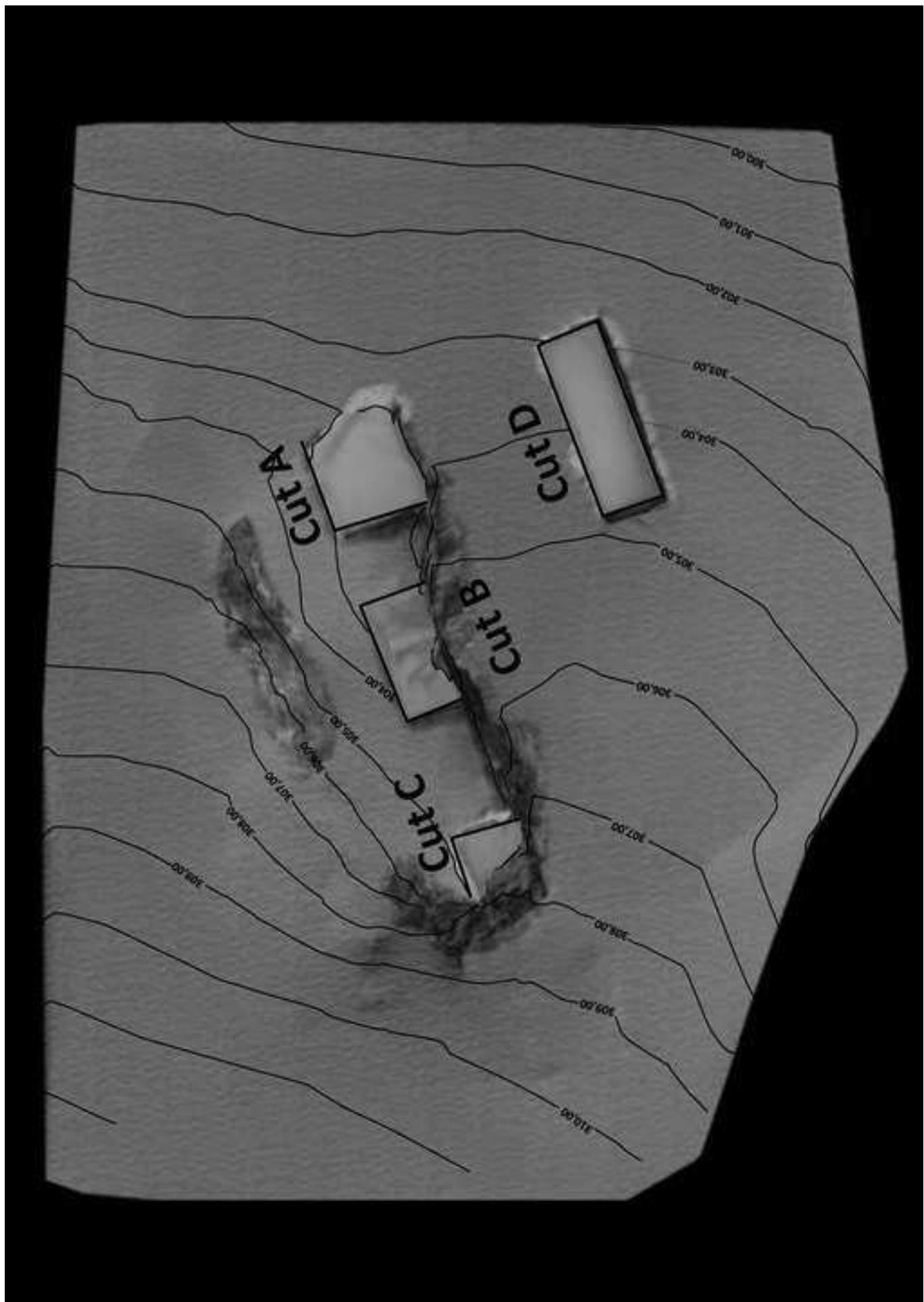


Figure
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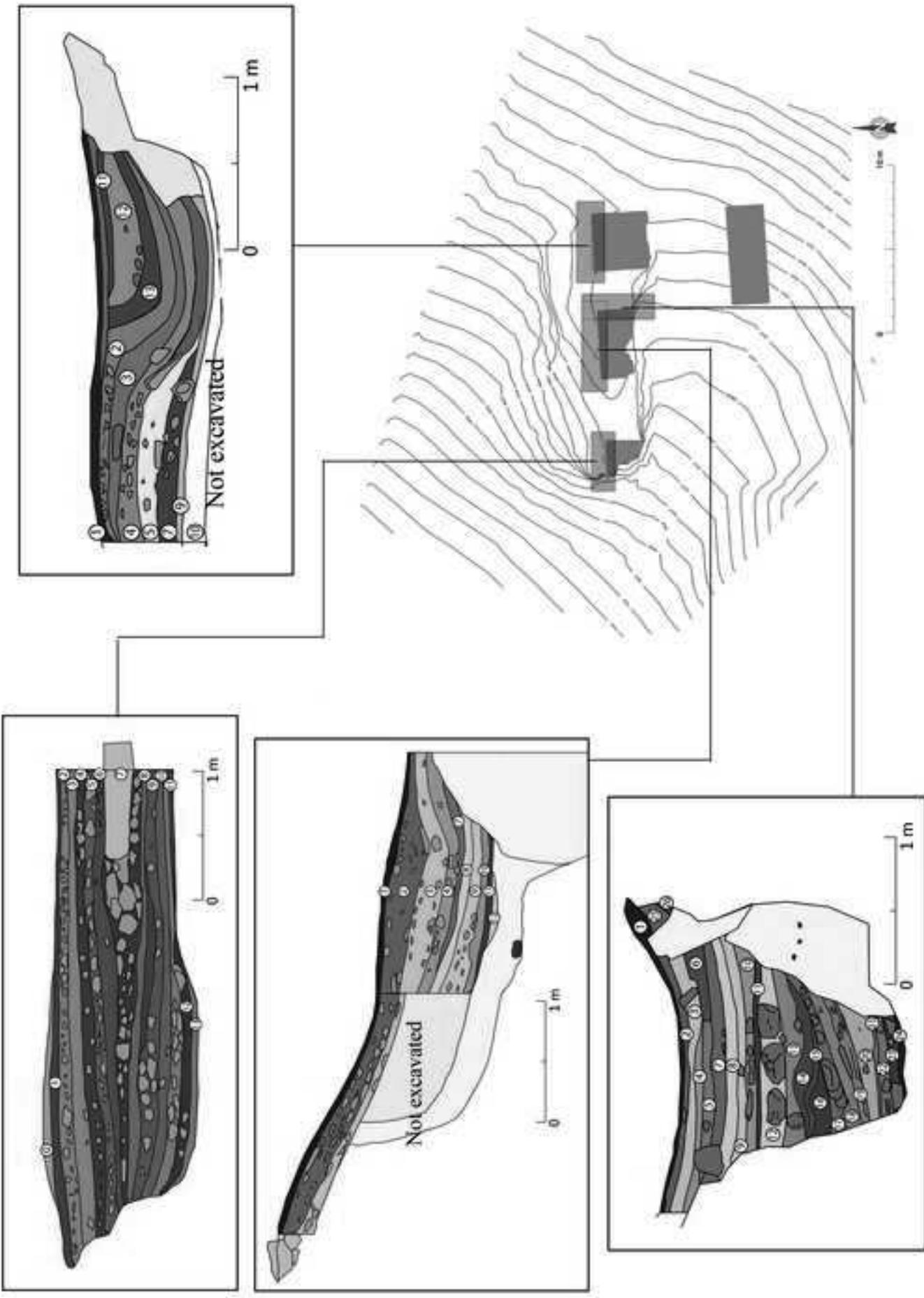
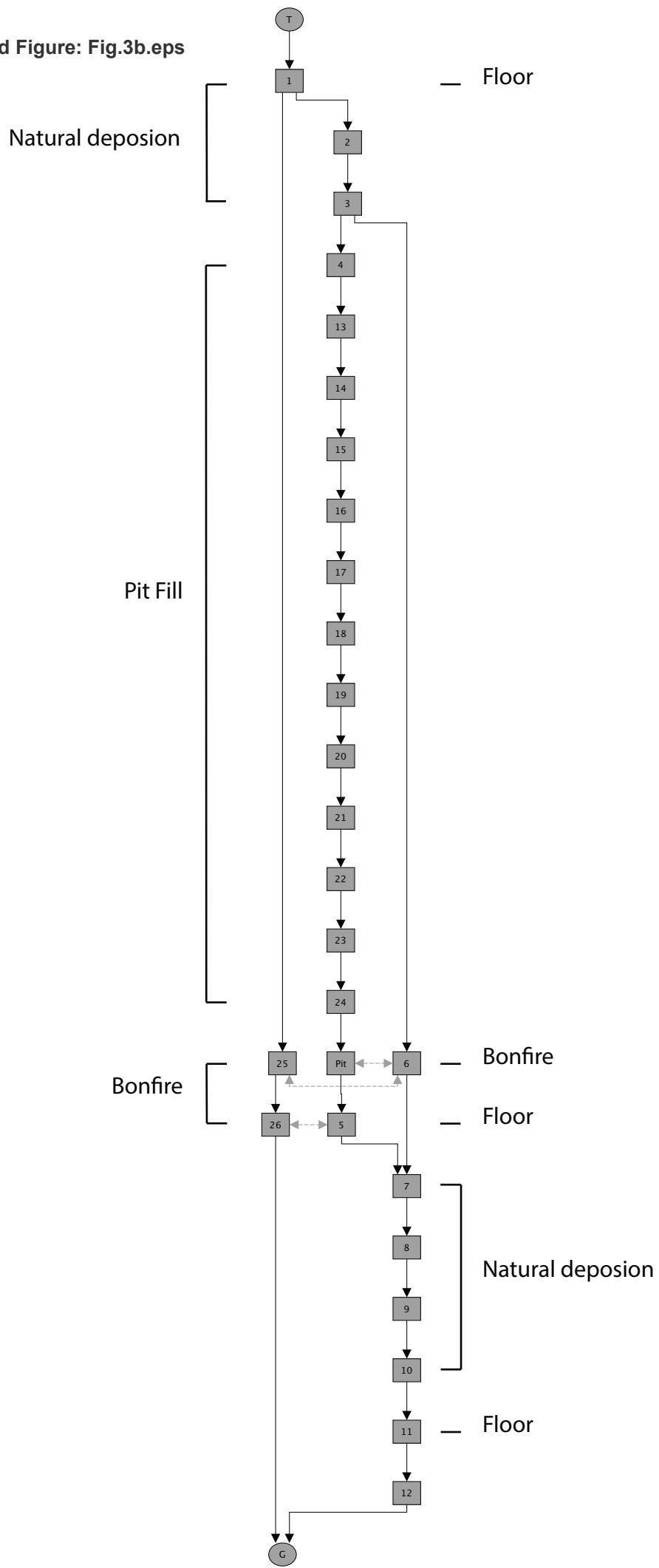


Figure3b

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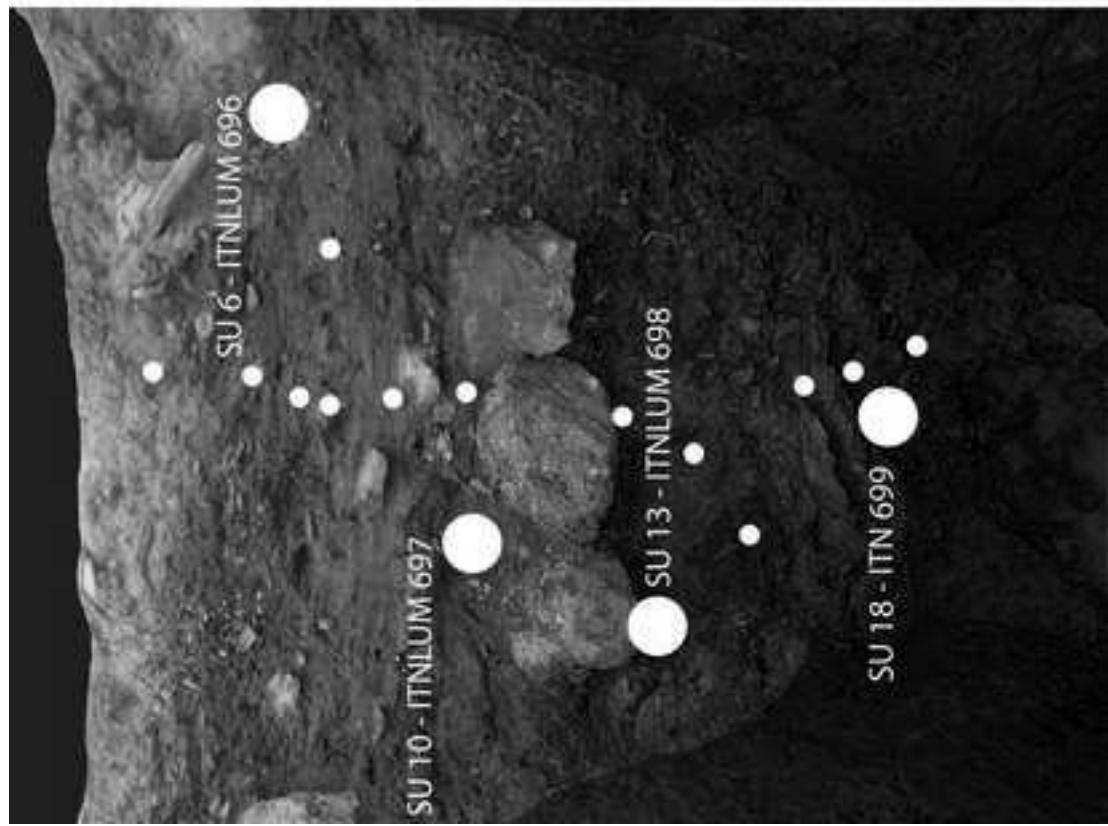
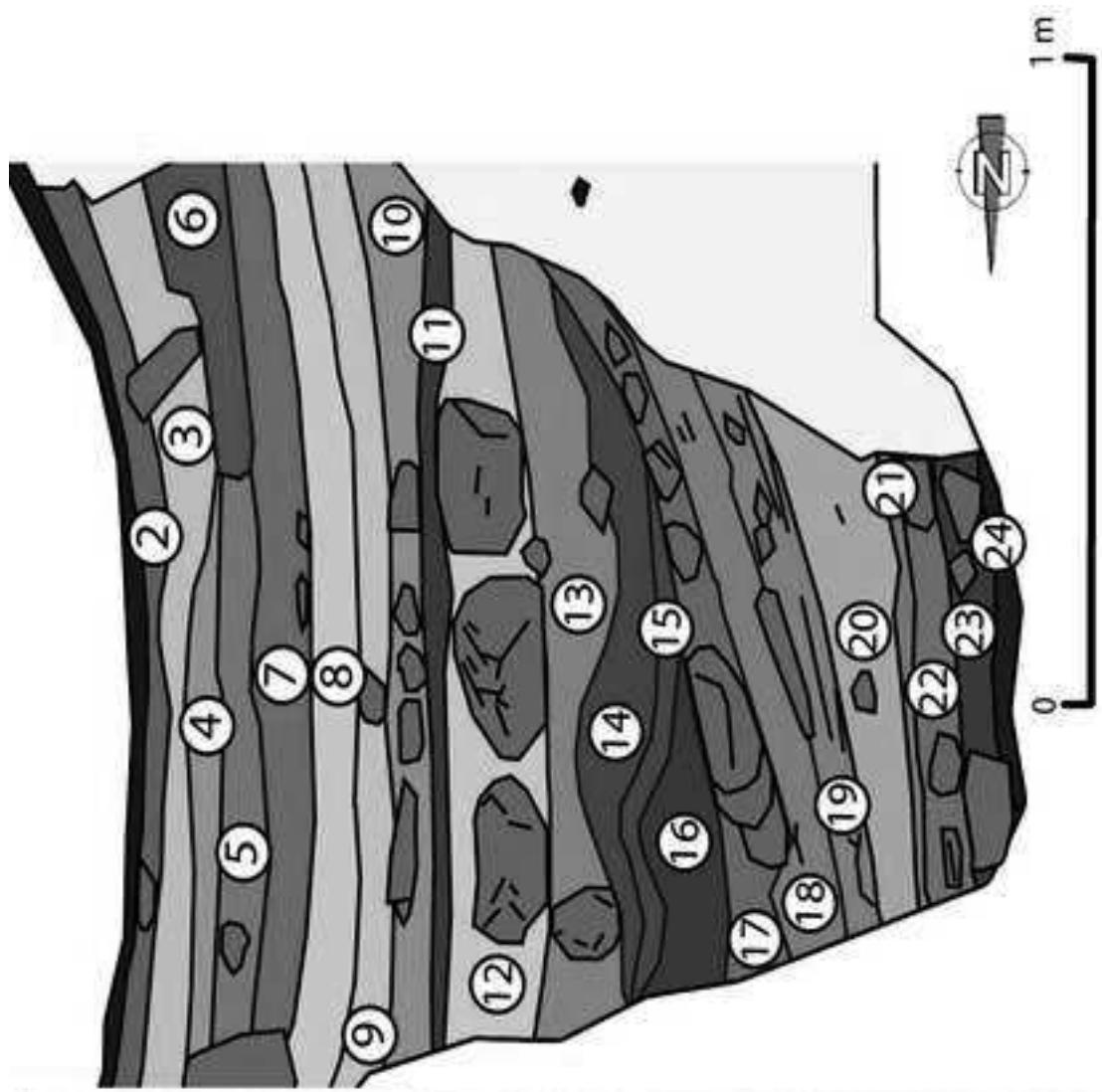


Figure4b

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**CUT D
NORTH PROFILE**

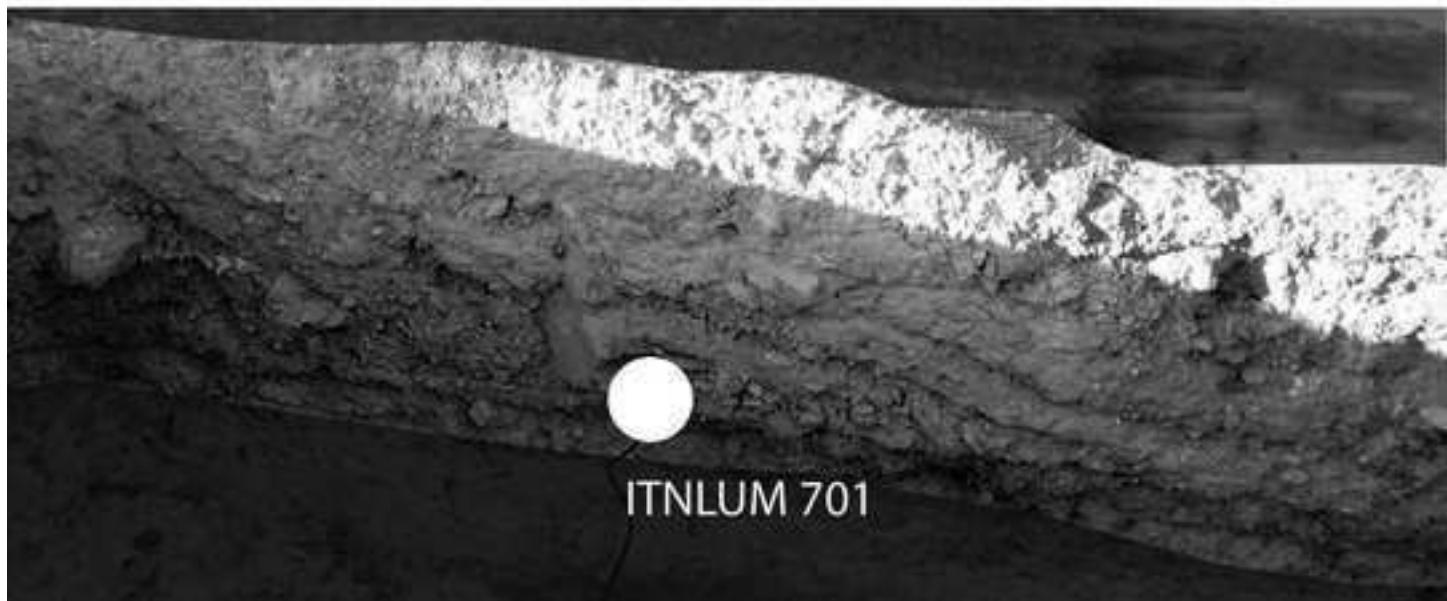
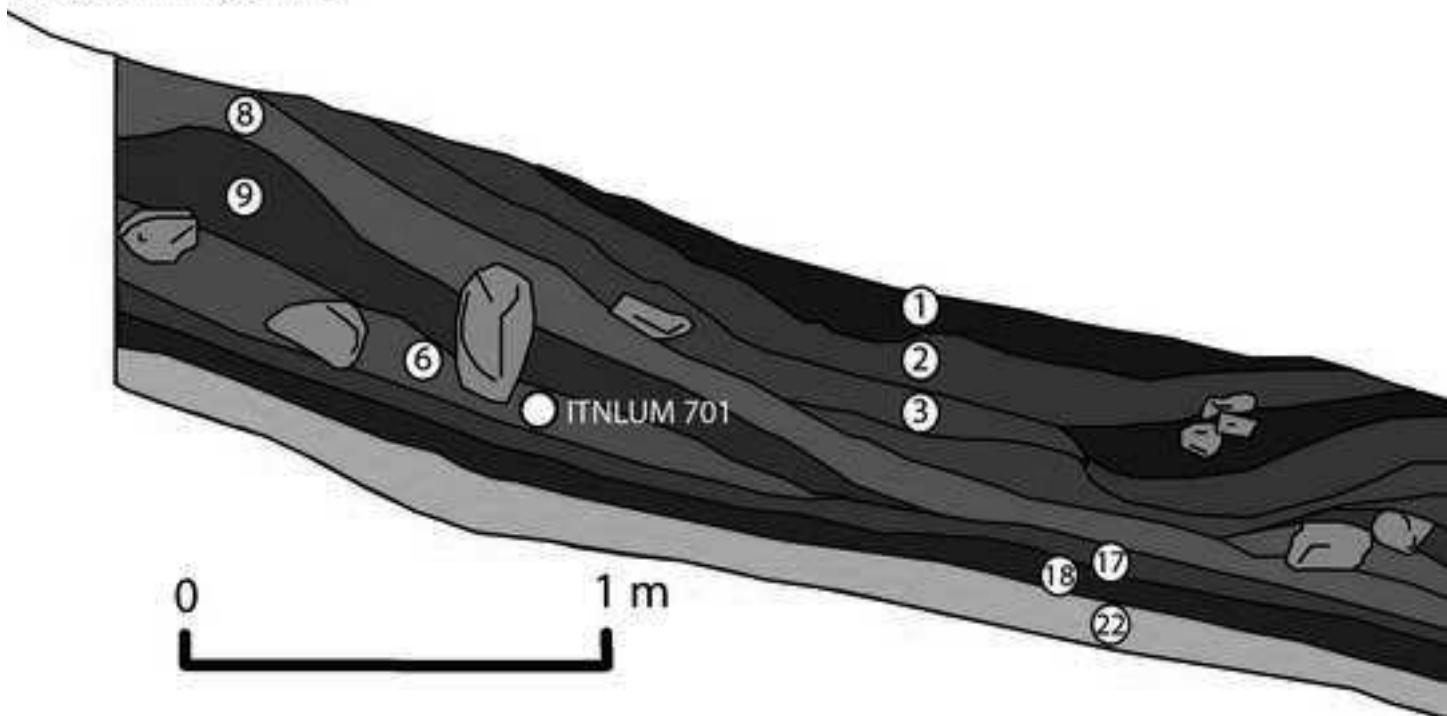


Figure 5

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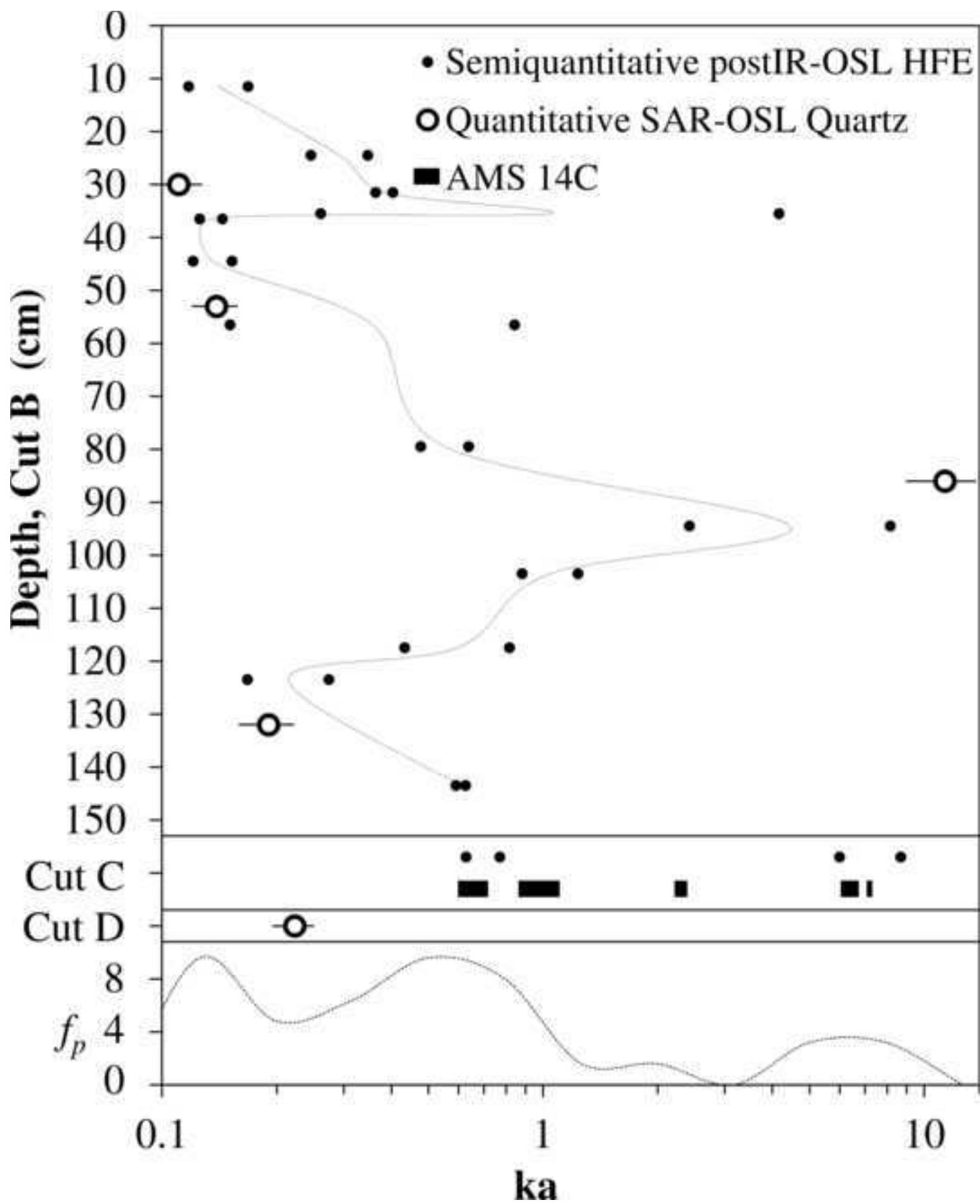


Figure6

TA DAS CASTELHANAS

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OUS A NOVA

ENTEPECINA II (SP)

LMEN DE ALBERITE (SP)

LE DE RODRIGO (SP)

RAPITO 1 (SP)

GUIJO

RMOTA

TA DA HORTA

SULLO

PA DO BUGIO (SP)

SA DA MORA (SP)

BEÇO DE ARRUDA 1 (SP)

IMOGO (SP)

PIJOTTILLA TOMB 3 (SP)

DRA BRANCA (SP)

DEAGORDILLO (SP)

PA DA RAINHA

INTA DO ANJO ART. CAVE 3 (SP)

CO VELHO (SP)

TOMILLAR

TA GRANDE DE ZAMBUJEIRO (corridor)

VA DA MOURA (SP)

RDIGÕES TOMB 1 & 2 (SP)

mprob

CEIA (SP)

S ITUEROS (SP)

O PAULO II (SP)

mprob

O PEDRO DE ESTORIL 1 (SP)

S GABRIELES (SP)

OLOS BORRACHEIRA (SP)

LMEN PUERTO DE LOS HUERTOS (SP)

ITA DA LADRA (SP)

NHA VERDE (SP)

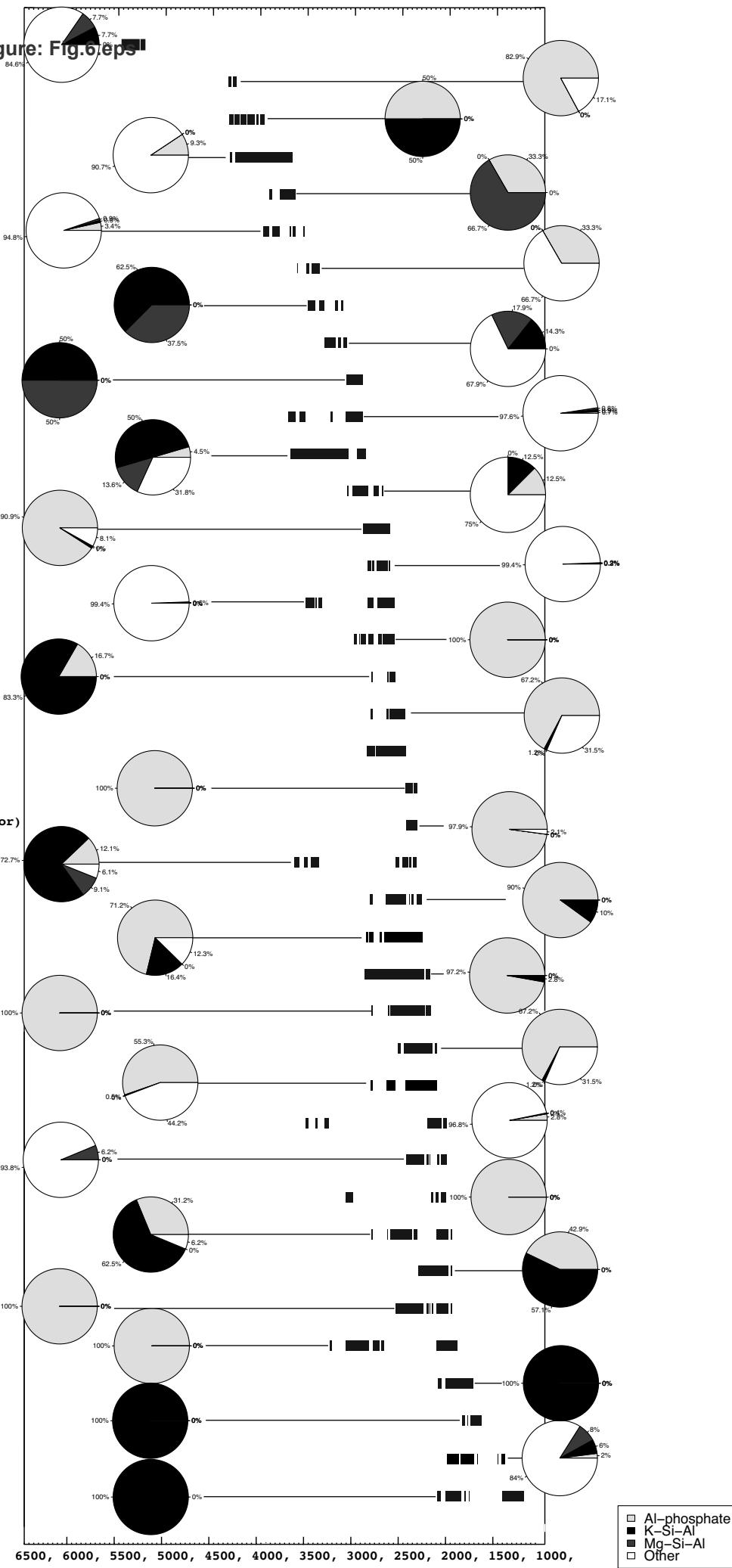
AIA MAÇAS (SP)

POZUELO (tomb 6)

S MINITAS (tomb 15)

ENTE ÁLAMO (SP)

STILLO DE ALANGE, C/UMBRIA (SP)



cal BC

□ Al-phosphate
■ K-Si-Al
■ Mg-Si-Al
○ Other

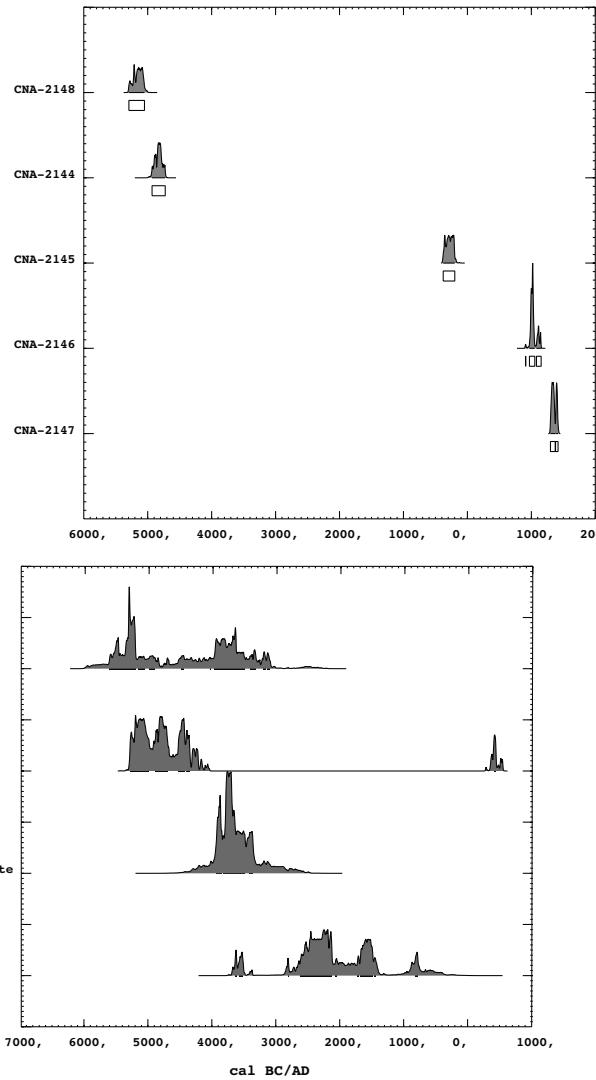
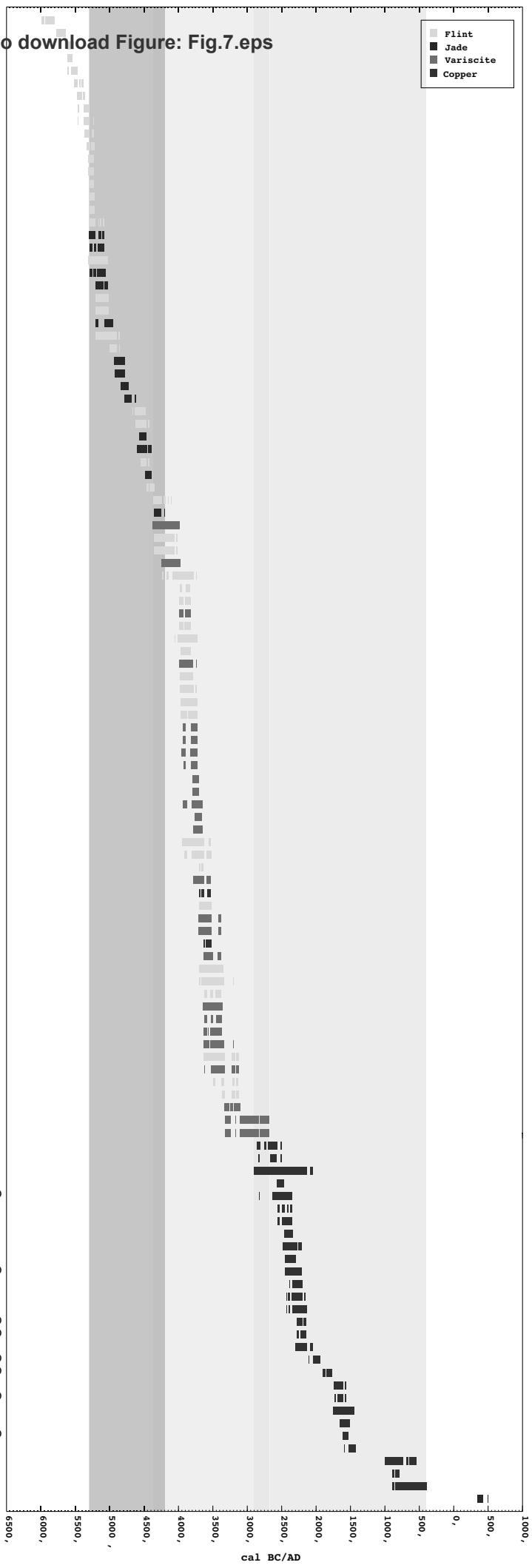
Figure7[Click here to download Figure: Fig.7.eps](#)

Table1

Table 1. Brief description of the OSL dating sampled units.

Sample #	Cut	Unit	Description
ITNLUM 696	B	6	A bonfire nearby the extraction surface.
ITNLUM 697	B	10	Rock sample. This unit seems to be directly related with the last episode of prehistoric exploitation. This unit accounts for the 15% of the stone tools recovered. Units is cut by the pit
ITNLUM 698	B	13	It is interpreted as a modern depositional unit filling the pit.
ITNLUM 699	B	18	Take part of the oblique depositional units that fills the pit with big rock blocks.
ITNLUM 701	D	6	Depositional unit, believed to be part of production debris accumulation.

Table 2a: Luminescence dating measurements: radionuclide concentrations and water content estimates.

PCMII	ITNLUM	Sample Type ¹	Depth ² (cm)	FGS <i>in situ</i> ³			H_2O <i>in situ</i> (g/g)	HRGS <i>lab, dry</i>			INAA <i>lab, dry</i>			H_2O time averaged (g/g)	
				K (%)	Th (ppm)	U (ppm)		Ref. A10/	K (%)	Th (ppm)	U (%)	Th (ppm)	U (ppm)		
1	696	T	30	1.5	9.1	7.6	0.37	256	2.4	13	16	2.2	12	14	0.10
2	697	T	53	1.7	9.6	7.9	0.36	257	2.3	14	15	2.3	12	14	0.19
3	698	R	86	1.5	8.7	7.6	0.37	258	1.8	11	13	1.9	11	11	0.19
4	699	T	132	1.8	7.3	14	0.38	259	2.6	14	23	2.6	12	23	0.20
5	701	T	47	1.7	11	8.1	0.08	260	1.8	12	14	1.7	12	13	0.09
<i>Average uncertainty</i>				0.1	1.2	1.1	0.02		0.1	1.8	0.6	0.2	0.7	0.4	0.15

¹ Tube; Rock. ² Below pre-excavation ground level. ³ As measured, i.e. not corrected for *in situ* water content.

Table 2b: Luminescence dating measurements: summary dose rate, dose, and calendar date estimates.

PCMII	ITNLUM	Sample Type ¹	Depth ² (cm)	Dose Rate		\bar{D}_{total}	σ_D	OSL Reader		Aliquots	Absorbed Dose		Calendar Date		
				\dot{D}_{cosmic}	σ_D			=	/24		\bar{I}_{T1}	D	σ_D	Date	σ_{date}
1	696	T	30	0.14	0.04	5.60	0.29	=	2	16	420	0.63	0.09	=	1900 20 AD
2	697	T	53	0.13	0.04	5.25	0.38	=	2	16	215	0.73	0.09	=	1870 20 AD
3	698	R	86	0.12	0.03	4.79	0.31	=	1	22	186	55	11	?	9000 2000 BC
4	699	T	132	0.12	0.03	6.97	0.59	=	3	16	362	1.34	0.19	>	1820 30 AD
5	701	T	47	0.22	0.01	4.79	0.27	=	3	16	2380	1.08	0.12	>	1790 30 AD

¹ Tube; Rock. ² Below pre-excavation ground level. ³ Summary assessment of dosimetry and luminescence results. =/=/>/? : equal to/overestimates/underestimates "true" value, within quoted uncertainties. ? : potentially outwith uncertainties. ⁴ Signal Intensity measured in response to the first Test dose in the SAR measurement sequence, averaged over all aliquots.

Table3

Table 3. Results of AMS-radiocarbon dating and anthracological analysis of sampled charcoals from PCM2.

Lab. Code	Specie	Sample	Context	Cut	Unit	Radiocarbon date	Calibrated date (2 σγμα)
CNA-2144	-	Charcoal	PCM2	C	12	5950±40	4033-4727 BC
CNA-2145	angiosperma	Charcoal	PCM2	C	12/13	2215±35	375-198 BC
CNA-2146	Cf. Cistus sp.	Charcoal	PCM2	C	12	1010±35	910-1152 AD
CNA-2147	Cf. Cistus sp., Cf. Pistacia lentiscus	Charcoal	PCM2	C	12	585±35	1298-1417 AD
CNA-2148	Cf. Quercus suber	Charcoal	PCM2	C	13	6205±40	5295-5051 BC

*According to the calibration curves IntCal13 (samples of the terrestrial biosphere) of Reimer et al. (2013), and using CALIB rev 7.1 (Stuiver and Reimer, 1993) program

Table4

Table 4. Results of AMS-radiocarbon dating and anthracological analysis of sampled charcoals from PCM2.

Lab. Code	Specie	Sample	Context	Cut	Unit	^{14}C yr BP	cal yr (2σ)
CNA-2144	-	Charcoal	PCM2	C	12	5950±40	4933-4727 cal BC
CNA-2145	angiosperma	Charcoal	PCM2	C	12/13	2215±35	379-198 cal BC
CNA-2146	Cf. <i>Cistus</i> sp.	Charcoal	PCM2	C	12	1010±35	969-1152 cal AD
CNA-2147	Cf. <i>Cistus</i> sp., Cf. <i>Pistacia</i> Charcoal lentiscus	Charcoal	PCM2	C	12	585±35	1298-1417 cal AD
CNA-2148	Cf. <i>Quercus</i> suber	Charcoal	PCM2	C	13	6205±40	5295-5051 cal BC

*According to the calibration curves IntCal13 (samples of the terrestrial biosphere) of Reimer et al. (2013), and using CALIB rev 7.1 (Stuiver and Reimer, 1993) program

Table 5. European mining resources available radiocarbon dates

Lab. code	Site	14C Age	14C Age	d13C	Mineral	Reference
Can Tintorer	I-12730	4310	150	0	variscita	(Bosch and Estrada, 1994)
Can Tintorer	I-12731	5350	190	0	variscita	(Bosch and Estrada, 1994)
Can Tintorer	I-11786	5070	100	0	variscita	(Bosch and Estrada, 1994)
Can Tintorer	UBAR-41	4970	100	0	variscita	(Bosch and Estrada, 1994)
Can Tintorer	CSIC-488	4710	50	0	variscita	(Bosch and Estrada, 1994)
Can Tintorer	CSIC-489	4940	50	0	variscita	(Bosch and Estrada, 1994)
Can Tintorer	I-12158	4880	100	0	variscita	(Bosch and Estrada, 1994)
Can Tintorer	UBAR-42	4820	100	0	variscita	(Bosch and Estrada, 1994)
Can Tintorer	I-13099	4820	100	0	variscita	(Bosch and Estrada, 1994)
Can Tintorer	UBAR-49	4740	90	0	variscita	(Bosch and Estrada, 1994)
Can Tintorer	UBAR-30	4710	130	0	variscita	(Bosch and Estrada, 1994)
Can Tintorer	UBAR-48	4690	100	0	variscita	(Bosch and Estrada, 1994)
Can Tintorer	UBAR-47	4610	90	0	variscita	(Bosch and Estrada, 1994)
Can Tintorer	I-12730	4310	150	0	variscita	(Bosch and Estrada, 1994)
Can Tintorer	Beta-61491	4660	110	0	variscita	(Bosch and Estrada, 1994)
Can Tintorer	Beta-72551	4930	70	0	variscita	(Bosch and Estrada, 1994)
Can Tintorer	Beta-72552	5000	60	0	variscita	(Bosch and Estrada, 1994)
Can Tintorer	Beta-72553	5100	60	0	variscita	(Bosch and Estrada, 1994)
S. Ferreteres	Beta-155686	5220	110	0	variscita	(Borrell et al., 2009)
S. Ferreteres	Beta-250402	5000	40	0	variscita	(Borrell et al., 2009)
S. Ferreteres	Beta-250403	4980	40	0	variscita	(Borrell et al., 2009)
S. Ferreteres	Beta-250405	4980	40	0	variscita	(Borrell et al., 2009)
S. Ferreteres	Beta-250406	5010	40	0	variscita	(Borrell et al., 2009)
S. Ferreteres	Beta-250404	5010	40	0	variscita	(Borrell et al., 2009)

Casa Montero	Beta-206512	6410	40	-24.2	silex	(Díaz del Rio and Consuegra Rodríguez, 2011)
Casa Montero	Beta-206513	6270	40	-26.2	silex	(Díaz del Rio and Consuegra Rodríguez, 2011)
Casa Montero	Beta-232884	6360	40	-25.4	silex	(Díaz del Rio and Consuegra Rodríguez, 2011)
Casa Montero	Beta-232885	6280	40	-24.9	silex	(Díaz del Rio and Consuegra Rodríguez, 2011)
Casa Montero	Beta-232886	6350	40	-25.6	silex	(Díaz del Rio and Consuegra Rodríguez, 2011)
Casa Montero	Beta-232887	6290	40	-22.2	silex	(Díaz del Rio and Consuegra Rodríguez, 2011)
Casa Montero	Beta-232888	6240	40	0	silex	(Díaz del Rio and Consuegra Rodríguez, 2011)
Casa Montero	Beta-232889	6290	40	-22.3	silex	(Díaz del Rio and Consuegra Rodríguez, 2011)
Casa Montero	Beta-232890	6500	40	-25.6	silex	(Díaz del Rio and Consuegra Rodríguez, 2011)
Casa Montero	Beta-232891	6320	40	-26.2	silex	(Díaz del Rio and Consuegra Rodríguez, 2011)
Casa Montero	Beta-232892	6270	40	-26.2	silex	(Díaz del Rio and Consuegra Rodríguez, 2011)
Casa Montero	Beta-232893	6330	40	-25.6	silex	(Díaz del Rio and Consuegra Rodríguez, 2011)
Defensola	UTC-1342	6990	80	0	Silex	(Díaz del Rio et al., 2008)
Defensola	Beta-71143	6820	80	0	Silex	(Díaz del Rio et al., 2008)
Defensola	Beta-71144	5670	70	0	Silex	(Díaz del Rio et al., 2008)
Defensola	Beta-80604	6630	70	0	Silex	(Díaz del Rio et al., 2008)
Defensola	Beta-80603	6540	60	0	Silex	(Díaz del Rio et al., 2008)
Defensola	UTC-1411	6630	40	0	Silex	(Díaz del Rio et al., 2008)
Tomaszów	GrN-7594	6145	70	0	Silex	(Díaz del Rio et al., 2008)
Krzemionki	Gd-1425	6090	110	0	Silex	(Díaz del Rio et al., 2008)
Porco	AA-62119	5665	47	0	jadeita	(Petrequin et al., 2006)
Porco	AA-62120	5847	47	0	jadeita	(Petrequin et al., 2006)
Porco	AA-62121	6110	48	0	jadeita	(Petrequin et al., 2006)
Porco	AA-62123	5959	49	0	jadeita	(Petrequin et al., 2006)
Porco	AA-62122	6146	49	0	jadeita	(Petrequin et al., 2006)
Porco	AA-62125	5931	48	0	jadeita	(Petrequin et al., 2006)
Porco	AA-62124	6231	48	0	jadeita	(Petrequin et al., 2006)
Bule	AA-66511	1644	36	0	jadeita	(Petrequin et al., 2006)

Bule	AA-66512	5393	42	0	jadeita	(Petrequin et al., 2006)
Bule	AA-66513	5605	42	0	jadeita	(Petrequin et al., 2006)
Bule	AA-66514	5662	71	0	jadeita	(Petrequin et al., 2006)
Bule	AA-66515	6222	44	0	jadeita	(Petrequin et al., 2006)
Bule	AA-66516	5963	61	0	jadeita	(Petrequin et al., 2006)
Bule	AA-66517	6212	71	0	jadeita	(Petrequin et al., 2006)
Cabrieres	Beta-156929	3830	40	0	cobre	(Amberg, 2002)
Cabrieres	Beta-156928	3900	40	0	cobre	(Amberg, 2002)
Araico	Beta-312351	5640	40	0	Silex	(Tarrío et al., 2011)
Araico	Beta-312352	6050	40	0	Silex	(Tarrío et al., 2011)
Tomaszów		6220	120	0	Silex	(Tarrío et al., 2011)
Tomaszów		6145	70	0	Silex	(Tarrío et al., 2011)
Blackpatch	BM-290	5090	130	0	silex	(Tarrío et al., 2011)
Church Hill	BM-181	5340	150	0	silex	(Whittle et al., 2011)
Cissbury	BM-183	4720	150	0	silex	(Whittle et al., 2011)
Cissbury	BM-184	4650	150	0	silex	(Whittle et al., 2011)
Cissbury	BM-185	4730	150	0	silex	(Whittle et al., 2011)
Cissbury	BM-3082	5100	60	-19.2	silex	(Whittle et al., 2011)
Cissbury	BM-3086	4710	60	-22.1	silex	(Whittle et al., 2011)
Harrow Hill	BM-182	4930	150	0	silex	(Whittle et al., 2011)
Harrow Hill	BM-2099R	5040	120	-23.1	silex	(Whittle et al., 2011)
Harrow Hill	BM-2097R	5140	150	-25.2	silex	(Whittle et al., 2011)
Harrow Hill	BM-2071R	4900	120	-26.7	silex	(Whittle et al., 2011)
Harrow Hill	BM-2075R	5020	110	-26.4	silex	(Whittle et al., 2011)
Harrow Hill	BM-2124R	5060	90	-24.9	silex	(Whittle et al., 2011)
Harrow Hill	BM-2098R	5350	150	-25.7	silex	(Whittle et al., 2011)
Harrow Hill	BM-3084	4880	30	-21.9	silex	(Whittle et al., 2011)

Harrow Hill	BM-3085	5070	50	-23.4	silex	(Whittle et al., 2011)
Long Down	OxA-1152	5050	100	0	silex	(Whittle et al., 2011)
Spiennes	Lv-1566	5510	55	0	silex	(Whittle et al., 2011)
Spiennes	GrN-4674	5420	75	0	silex	(Collet, 2004; 2008)
Spiennes	Lv-1598	5100	65	0	silex	(Collet, 2004; 2008)
Spiennes	KN-I.16	5110	40	0	silex	(Collet, 2004; 2008)
Spiennes	OxA-3196	4830	80	0	silex	(Collet, 2004; 2008)
Spiennes	Beta-194770	4580	40	0	silex	(Collet, 2004; 2008)
Spiennes	Beta-194771	4550	40	0	silex	(Collet, 2004; 2008)
Spiennes	Beta-110683	4500	50	0	silex	(Collet, 2004; 2008)
El Milagro (Asturias)	OxA-3005	3990	90	0	cobre	(Collet, 2004; 2008)
El Milagro (Asturias)	OxA-3006	3850	90	0	cobre	(Blas Cortina, 2011)
El Milagro (Asturias)	Ua-33207	3785	35	0	cobre	(Blas Cortina, 2011)
El Milagro (Asturias)	Ua-33209	3775	35	0	cobre	(Blas Cortina, 2011)
El Milagro (Asturias)	Ua-24538	3630	40	0	cobre	(Blas Cortina, 2011)
El Milagro (Asturias)	Ua-24537	3520	40	0	cobre	(Blas Cortina, 2011)
El Milagro (Asturias)	Ua-24550	3355	45	0	cobre	(Blas Cortina, 2011)
El Milagro (Asturias)	Ua-33206	3285	35	0	cobre	(Blas Cortina, 2011)
La Profunda (León)	Ua-35778	3865	35	0	cobre	(Blas Cortina, 2011)
La Profunda (León)	Ua-35779	3950	35	0	cobre	(Blas Cortina, 2011)
La Profunda (León)	Ua-35780	4075	35	0	cobre	(Blas Cortina, 2011)
El Aramo (Asturias)	OxA-1833	4090	70	0	cobre	(Blas Cortina, 2011)
El Aramo (Asturias)	OxA-1926	3810	70	0	cobre	(Blas Cortina, 2011)
El Aramo (Asturias)	OxA-3007	3900	90	0	cobre	(Blas Cortina, 2011)
El Aramo (Asturias)	OxA-6789	3995	50	0	cobre	(Blas Cortina, 2011)
El Aramo (Asturias)	Ua-18629	3775	65	-21.4	cobre	(Blas Cortina, 2011)
El Aramo (Asturias)	Ua-18630	3365	60	-20.1	cobre	(Blas Cortina, 2011)
El Aramo (Asturias)	Ua-18631	3310	65	-20.4	cobre	(Blas Cortina, 2011)

El Aramo (Asturias)	Ua-18632	3825	60	-20.5	cobre	(Blas Cortina, 2011)
El Aramo (Asturias)	Ua-18633	3940	60	-20.1	cobre	(Blas Cortina, 2011)
El Aramo (Asturias)	Ua-18634	3215	55	-22	cobre	(Blas Cortina, 2011)
Chiflón	BM-1600	4840	50	0	cobre	(Blas Cortina, 2011)
Chiflón	BM-1599	4780	50	0	cobre	(Acosta, 1995)
Chiflón	OxTL-200e3(II)	4000	300	0	cobre	(Acosta, 1995)
Rio Tinto	BM-2337R	2650	140	0	cobre	(Acosta, 1995)
Chiflón	BM-1529	3320	130	0	cobre	(Castro Martínez et al., 1996)
Chiflón	BM-1528	2650	60	0	cobre	(Burleigh et al., 1982)
Chiflón	BM-1601	2520	210	0	cobre	(Burleigh et al., 1982; Rothenberg and Freijeiro, 1980)

Table 6: Studied sites available radiocarbon dates

Lab. code	Site	14C Age	14C Age	d13C	Reference
Beta-83084 GrN-19168	ALDEAGORDILLO	4320 years BP	70 years SD	per mil 0	(Fabián García, 2006)
Beta-83085	ALDEAGORDILLO	4115	20	0	(Fabián García, 2006)
Beta-194313	ALDEAGORDILLO	4100	80	0	(Fabián García, 2006)
ICEN-1264	ANTA DA HORTA	4480	40	-19.7	(Oliveira, 2010)
Beta-243693	ANTA DAS CASTELHANAS	6360	110	0	(Oliveira, 2000)
Beta-123363	ANTA GRANDE DE ZAMBUJEIRO (corridor)	3910	40	0	(Soares and Silva, 2010)
Beta-132975	CABEÇO DA ARRUDA I	4370	70	0	(Waterman, 2012)
GrN-5110	CABEÇO DA ARRUDA I	4240	50	0	(Waterman, 2012)
GrN-5110	CARAPITO 1	4850	40	0	(Senna-Martínez and Quintá Ventura, 2000)
OxA-3733	CARAPITO 1	5125	70	0	(Senna-Martínez and Quintá Ventura, 2000)
TO-3336	CARAPITO 1	5120	40	0	(Senna-Martínez and Quintá Ventura, 2000)
OxA-5506 TO-953	CASA DA MOURA	4600	90	0	(Cardoso et al., 1996)
TO-2092	CASA DA MOURA	5990	60	-19.6	(Cardoso et al., 1996)
TO-2093	CASA DA MOURA	4850	100	-19.3	(Cardoso et al., 1996)
TO-2094	CASA DA MOURA	5070	70	-19.2	(Cardoso et al., 1996)
ICEN-802	CASA DA MOURA (1A)	5020	70	-19.6	(Cardoso et al., 1996)
Beta-68667	CASTILLO DE ALANGE (c/ Umbría 3/ N. II)	6100	70	0	(Cardoso et al., 1996)
CNA-346	CASULLO	3080	90	0	(Pavón Soldevila, 1994)
Beta-277240 UBAR-593	CHOUZA NOVA	4410	50	0	(Linares Catela and García Sanjuán, 2010)
UBAR-536	COVA DA MOURA	5450	40	0	(Dominguez Bella and Bóveda, 2011)
ICEN-1040	COVA DA MOURA	4715	50	0	(Silva, 2003)
ICEN-1041	DOLMEN DA PEDRA BRANCA (chamber)	3950	60	0	(Silva, 2003)
	DOLMEN DA PEDRA BRANCA (corridor)	4620	60	-19.7	(Soares, 2010)
		4120	60	-	(Soares, 2010)
Beta-80602 Beta-80600	DOLMEN DE ALBERITE	5320	70	20.06	(Stipp and Tamers, 1996)
Beta-80598	DOLMEN DE ALBERITE	5110	140	0	(Stipp and Tamers, 1996)
Poz-55021	DOLMEN DE ALBERITE	5020	70	0	(Stipp and Tamers, 1996)
Teledyne 19080	EL GUIJO	4695	35	0	(Villalobos García, 2014)
GrN-18875	EL POZUELO (tomb 6)	3580	120	0	(Linares Catela and García Sanjuán, 2010)
GrN-18669	EL TOMILLAR	3925	40	0	(Fabián García, 2006)
GrN-16073	FUENTEPEPECINA II	5375	45	0	(Delibes de Castro and Rojo Guerra, 1997)
	FUENTEPEPECINA II	100	0	0	(Delibes de Castro and Rojo Guerra, 1997)

Beta-176899	HERDADE DOS CEBOLINHOS (ANTA 2/ chamber)	3900	40	0	(Gonçalves, 2003)
Beta-177471	HERDADE DOS CEBOLINHOS (ANTA 2)	3840	40	0	(Gonçalves, 2003)
Beta-121143	LA PIJOTILLA (T3 UE 15)	4130	40	0	(Odriozola et al., 2008)
ITN A6/209	LA PIJOTILLA (T3) layer 16/ vessel 15)	2716	96	0	(Odriozola et al., 2008)
CNA-034	LA PIJOTILLA (T3) layer 18)	4168	55	0	(Odriozola et al., 2008)
UGAMS 8455	LAPA DA RAINHA	4080	25	0	(Waterman, 2012)
GrN-5628	LAPA DO BUGIO	4850	45	0	(Silva and Wasterlain, 2010)
OxA-5507	LAPA DO BUGIO	4420	110	0	(Silva and Wasterlain, 2010)
Beta-142035	LAS MINITAS (tomb 15)	3430	50	0	(Pavón Soldevila, 1994)
ICEN-674	LECEIA	4370	60	0	(Cardoso, 2014)
Wk-34424	LECEIA	3833	26	0	(Cardoso, 2014)
ICEN-95	LECEIA (C2)	4370	60	1.34	(Cardoso, 2014)
ICEN-89	LECEIA (C2)	4200	70	-	(Cardoso, 2014)
ICEN-92	LECEIA (C2)	4120	80	19.91	(Cardoso, 2014)
ICEN-102	LECEIA (C2)	3970	70	24.56	(Cardoso, 2014)
ICEN-1217	LECEIA (C2)	4020	80	1.68	(Cardoso, 2014)
ICEN-1220	LECEIA (C2)	4030	70	-	(Cardoso, 2014)
ICEN-737	LECEIA (C2)	3920	70	20.05	(Cardoso, 2014)
ICEN-315	LECEIA (C2)	3730	170	-	(Cardoso, 2014)
ICEN-1213	LECEIA (C2)	3970	70	19.56	(Cardoso, 2014)
ICEN-1218	LECEIA (C2)	3910	60	21.19	(Cardoso, 2014)
ICEN-1211	LECEIA (C2)	3900	80	-	(Cardoso, 2014)
ICEN-1215	LECEIA (C2)	3900	70	23.21	(Cardoso, 2014)
ICEN-1216	LECEIA (C2)	3880	80	-	(Cardoso, 2014)
ICEN-1214	LECEIA (C2)	3840	110	-	(Cardoso, 2014)
ICEN-314	LECEIA (C2)	3770	130	-	(Cardoso, 2014)
ICEN-91	LECEIA (C3)	4130	60	-20	(Cardoso, 2014)
ICEN-673	LECEIA (C3)	4130	100	-24.9	(Cardoso, 2014)

ICEN-675	LECEIA (C3)	4100	90	-25.4	(Cardoso, 2014)
ICEN-1173	LECEIA (C3)	4170	50	-20.5	(Cardoso, 2014)
ICEN-1175	LECEIA (C3)	4090	80	-	(Cardoso, 2014)
ICEN-1176	LECEIA (C3)	4090	60	19.85	(Cardoso, 2014)
ICEN-1177	LECEIA (C3)	4050	50	20.02	(Cardoso, 2014)
LY-4205	LECEIA (C3)	4030	120	-21.2	(Cardoso, 2014)
ICEN-1160	LECEIA (C4)	4630	60	0	(Cardoso, 2014)
ICEN-312	LECEIA (C4)	4530	100	0	(Cardoso, 2014)
ICEN-316	LECEIA (C4)	4520	70	0	(Cardoso, 2014)
ICEN-312	LECEIA (C4)	4530	100	-	(Cardoso, 2014)
ICEN-313	LECEIA (C4)	4520	130	20.22	(Cardoso, 2014)
ICEN-316	LECEIA (C4)	4520	70	-22	(Cardoso, 2014)
ICEN-1275	LECEIA (EN)	3950	90	0	(Cardoso, 2014)
ICEN-1241	LECEIA (EN)	3950	90	0	(Cardoso, 2014)
Beta-260295	LECEIA (EN)	3840	40	0	(Cardoso, 2014)
Beta-260296	LECEIA (EN)	3980	40	0	(Cardoso, 2014)
Sac-1317	LECEIA (FM)	4220	50	0	(Cardoso, 2014)
Sac-1317	LECEIA (FM)	4220	50	0	(Cardoso, 2014)
Beta-260297	LECEIA (FM)	4140	40	0	(Cardoso, 2014)
Beta-260299	LECEIA (FM)	4100	40	0	(Cardoso, 2014)
Sac-1317	LECEIA (FM)	4220	40	0	(Cardoso, 2014)
Beta-260297	LECEIA (FM)	4140	40	0	(Cardoso, 2014)
Beta-260299	LECEIA (FM)	4100	40	0	(Cardoso, 2014)
ICEN-1160	LECEIA (layer 4)	4630	45	-	(Cardoso, 2014)
ICEN-1161	LECEIA (layer 4)	4440	50	21.81	(Cardoso, 2014)
ICEN-1159	LECEIA (layer 4)	4430	50	23.39	(Cardoso, 2014)
ICEN-1158	LECEIA (layer 4)	4320	50	21.35	(Cardoso, 2014)
Beta-185650	LOS GABRIELES (dolmen 4/ chamber 2)	3700	50	21.45	(Linares Catela and García Sanjuán, 2010)
Beta-185648	LOS GABRIELES (dolmen 4/ chamber 3)	3850	40	0	(Linares Catela and García Sanjuán, 2010)
Beta-185649	LOS GABRIELES (dolmen 4/ chamber 4)	3920	50	0	(Linares Catela and García Sanjuán, 2010)
I-16150	LOS ITUEROS	4120	130	0	(Fabrián García, 2006)
I-83088	LOS ITUEROS	3960	90	0	(Fabrián García, 2006)

I-16149	100	0	(Fabián García, 2006)
OxA-5535	55	0	(Gonçalves, 2002)
Sac-2122	50	0	(Cardoso and Caninas, 2010)
Sac-2123	50	0	(Cardoso and Caninas, 2010)
Sac-2370	80	0	(Cardoso and Caninas, 2010)
Sac-2371	60	0	(Cardoso and Caninas, 2010)
ICEN-957	60	0	(Gonçalves, 2002)
ICEN-956	80	0	(Gonçalves, 2002)
ICEN-955	100	0	(Gonçalves, 2002)
Sac-1556	90	0	(Gonçalves, 2002)
UBAR-539	90	0	(Gonçalves, 2002)
Beta-186854	50	-20.5	(Sousa, 2010)
Beta-142451	40	0	(Sousa, 2010)
Sac-267	50	-	(Sousa, 2010)
Sac-2156	3640	0	(Sousa, 2010)
LOS ITUEROS	9)		
MARMOTA 2	3760	40	-19.6 (Sousa, 2010)
MOITA DA LADRA	4080	40	-21.2 (Sousa, 2010)
MOITA DA LADRA	3880	60	0 (Sousa, 2010)
MOITA DA LADRA	3930	30	-21.2 (Sousa, 2010)
MOITA DA LADRA	3760	50	0 (Sousa, 2010)
OLIVAL DA PEGA 2B (phase VF1/OP4)	4100	40	-20.2 (Sousa, 2010)
OLIVAL DA PEGA 2B (phase VF3/OP2)	3420	200	0 (Cardoso, 2010)
OLIVAL DA PEGA 2B (phase VF3/OP3)	3420	200	0 (Cardoso, 2010)
PAIMOGO	3420	200	0 (Cardoso, 2010)
PENEDO DE LEXIM (locus 1/ layer 19)	4000	40	0 (Cardoso, 2010)
PENEDO DE LEXIM (locus 1/ layer 19)	3970	40	0 (Cardoso, 2010)
PENEDO DE LEXIM (locus 1/ layer 19)	3890	40	0 (Cardoso, 2010)
PENEDO DE LEXIM (locus 3/ layer 19)	3700	30	0 (Cardoso, 2010)
PENEDO DE LEXIM (locus 3b/ layer 10)	3680	40	0 (Cardoso, 2010)
PENEDO DE LEXIM (locus 3b/ layer 7)	3760	50	0 (Cardoso, 2010)
PENEDO DE LEXIM (locus 3b/ layer 7b)	3930	30	-21.2 (Sousa, 2010)
PENEDO DE LEXIM (locus 5/ layer 8)	4100	40	-20.2 (Sousa, 2010)
PENEDO DE LEXIM (locus3b/ layer 16)	3420	200	0 (Cardoso, 2010)
PENHA VERDE	3420	200	0 (Cardoso, 2010)
PENHA VERDE (22/064)	4090	30	0 (Valera et al., 2014)
PENHA VERDE (ditch HUT 2)	4060	30	0 (Valera et al., 2014)
PENHA VERDE (HUT 1)	4130	30	0 (Valera et al., 2014)
PENHA VERDE (HUT 2)	3840	30	0 (Valera et al., 2014)
PENHA VERDE (HUT 3)	3830	40	0 (Cardoso, 2010)
PENHA VERDE (pavement HUT 2)	4030	40	0 (Valera et al., 2014)
PERDIGÕES (Tomb 1)	4090	30	0 (Valera et al., 2014)
PERDIGÕES (Tomb 1)	3890	30	0 (Valera et al., 2014)
PERDIGÕES (Tomb 1)	3970	30	0 (Valera et al., 2014)
PERDIGÕES (Tomb 2)	4245	55	0 (Gonçalves, 2008)
PERDIGÕES (Tomb 2)	4090	55	0 (Gonçalves, 2008)
POÇO VELHO	4090	40	-19.1 (Gonçalves, 2008)
POÇO VELHO	4090	40	-19.1 (Gonçalves, 2008)
POÇO VELHO (F-2)	4090	40	-19.1 (Gonçalves, 2008)

Beta-245137	4030	40	-10.2	(Gonçalves, 2008)
Beta-244397	3920	40	-19.8	(Gonçalves, 2008)
Beta-244394	4520	40	-19.1	(Gonçalves, 2008)
Beta-245138	4500	40	-19.1	(Gonçalves, 2008)
Beta-244395	4030	40	-18.5	(Gonçalves, 2008)
OxA-5533	4245	55	-19.4	(Gonçalves, 2008)
Beta-244393	4160	50	-19.1	(Gonçalves, 2008)
Beta-178464	4150	40	-19.3	(Gonçalves, 2008)
Beta-244390	4150	40	0	(Gonçalves, 2008)
OxA-5532	4090	55	-19.6	(Gonçalves, 2008)
Beta-244392	3970	40	-18.8	(Gonçalves, 2008)
Beta-178463	3960	40	-19.7	(Gonçalves, 2008)
Beta-244391	48090	1200	-20.1	(Gonçalves, 2008)
H-2049/148	4260	60	0	(Gonçalves, 2002)
H-2048/1458	3650	60	0	(Gonçalves, 2002)
OxA-5509	4410	75	0	(Gonçalves, 2002)
OxA-5510	4395	60	0	(Gonçalves, 2002)
Leisner & Freireira/ 1963	4160	110	0	(Gonçalves, 2002)
Leisner & Freireira/ 1963	3650	100	0	(Gonçalves, 2002)
CNA-342	4070	50	0	(Linares Catela and García Sanjuán, 2010)
CNA-344	3940	50	0	(Linares Catela and García Sanjuán, 2010)
CNA-341	3680	50	0	(Linares Catela and García Sanjuán, 2010)
GrN-10744	4040	70	0	(Gonçalves, 2002)
OxA-5508	4050	60	0	(Gonçalves, 2002)
UBAR-629	3960	190	0	(Silva, 2003)
UBAR-630	3870	70	0	(Silva, 2003)
Beta-188390	4720	40	-19	(Gonçalves, 2005)
Beta-178467	3830	40	-19.4	(Gonçalves, 2005)
Beta-178468	3790	40	-19.6	(Gonçalves, 2005)
UGAMS 8455	4420	25	0	(Waterman, 2012)
UGAMS 8454	3720	25	0	(Waterman, 2012)
Ua-10831	3905	75	0	(Boaventura, 2011)
Ua-10830	4905	60	0	(Boaventura, 2011)
KIA-3-1381	4996	29	0	(Boaventura, 2011)
KIA-27559	4238	29	0	(Kunst and Lutz, 2011)
KIA-7260	4134	43	0	(Kunst and Lutz, 2011)
KIA-27563	4065	37	0	(Kunst and Lutz, 2011)
KIA-28669	4001	28	0	(Kunst and Lutz, 2011)
KIA-27557	3996	23	0	(Kunst and Lutz, 2011)

KIA-27558	31	0	(Kunst and Lutz, 2011)
KIA-7259	3801	43	(Kunst and Lutz, 2011)
KIA-7258	3891	43	(Kunst and Lutz, 2011)
KIA-4509	3960	44	(Kunst and Lutz, 2011)
KIA-7256	3951	55	(Kunst and Lutz, 2011)
GrN-7009	4200	40	(Kunst and Lutz, 2011)
KIA-7257	3836	39	(Kunst and Lutz, 2011)
GrN-6671	4170	55	(Kunst and Lutz, 2011)
GrN-7002	4050	40	(Kunst and Lutz, 2011)
KIA-27561	4155	32	(Kunst and Lutz, 2011)
KIA-27562	4049	25	(Kunst and Lutz, 2011)
KN-4989	3917	50	(Kunst and Lutz, 2011)
KN-4990	3934	51	(Kunst and Lutz, 2011)
KN-4988	3980	40	(Kunst and Lutz, 2011)
KIA-7261	3842	37	(Kunst and Lutz, 2011)
KIA-27564	3992	32	(Kunst and Lutz, 2011)
KN-I.115	3530	65	(Kunst and Lutz, 2011)
GrN-7003	4055	40	(Kunst and Lutz, 2011)
GrN-7004	3995	35	(Kunst and Lutz, 2011)
GrN-7005	4055	40	(Kunst and Lutz, 2011)
GrN-7008	3980	35	(Kunst and Lutz, 2011)
GrN-6670	4150	105	(Kunst and Lutz, 2011)
KIA-28668	3999	29	(Kunst and Lutz, 2011)
GrN-7006	4090	40	(Kunst and Lutz, 2011)
GrN-6669	4025	95	(Kunst and Lutz, 2011)
GrN-7007C	3950	65	(Kunst and Lutz, 2011)
GrN-6668	3625	65	(Kunst and Lutz, 2011)
KN-4507	3466	53	(Kunst and Lutz, 2011)
KN-4506	3847	34	(Kunst and Lutz, 2011)
KIA-27565	4445	31	(Kunst and Lutz, 2011)
KIA-27556	3965	32	(Kunst and Lutz, 2011)
KIA-27566	3467	36	(Kunst and Lutz, 2011)
KIA-27641	2381	40	(Kunst and Lutz, 2011)

/video

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